

Problem set 8

Problem 1

- a) A quark confined to a region of linear dimension R has by Heisenberg's uncertainty principle a momentum

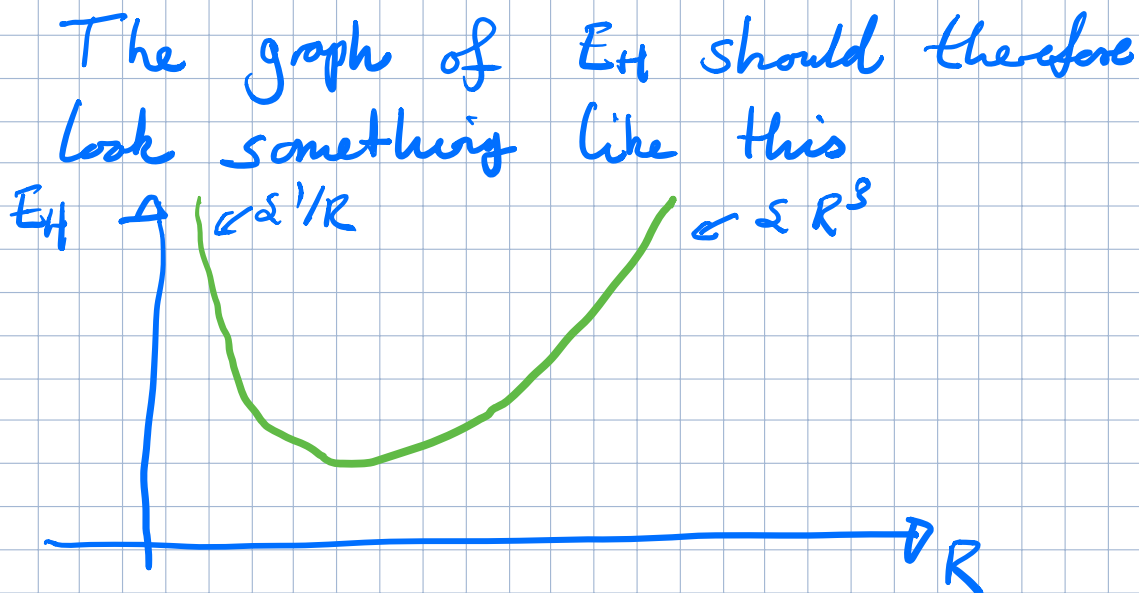
$$p \sim \frac{h}{R}$$

For a free, relativistic quark the energy is then

$$E = pc \sim \frac{hc}{R} \propto \frac{1}{R}$$

b)
$$E_H = \frac{4\pi}{3} BR^3 + \frac{C}{R}$$

For small R , the last term dominates, and $E_H \rightarrow \infty$ as $R \rightarrow 0$ (assuming C is positive, which is reasonable, since this term represents kinetic energy.) For large R , the first term dominates, and $E_H \rightarrow \infty$ as $R \rightarrow \infty$.



E_H should therefore have a minimum, and at this minimum we have

$$\frac{dE_H}{dR} = 0$$

$$\Rightarrow 4\pi B R^2 - \frac{C}{R^2} = 0$$

$$\Rightarrow \underline{\underline{C = 4\pi B R^4}}$$

Inserting this in E_H gives

$$\begin{aligned} E_{H, \text{min}} &= \frac{4\pi}{3} B R^3 + 4\pi B R^3 \\ &= \underline{\underline{\frac{16\pi}{3} B R^3}} \end{aligned}$$

c) With $E_{H, \min} = 10^3 \text{ MeV}$ ($\approx m_n c^2, m_p c^2$)
and $R = 10^{-15} \text{ m} = 1 \text{ fm}$

$$B = \frac{3}{16\pi} \frac{E_{H, \min}}{R^3} = \frac{3}{16\pi} \frac{10^3 \text{ MeV}}{1 \text{ fm}^3}$$
$$= \underline{\underline{59,7 \text{ MeV fm}^{-3}}}$$

d) We want $B = 200^4 \text{ MeV}^4$
in natural units. Since
 $\hbar c$ has units $\text{MeV} \cdot \text{fm}$,
we can translate this to
"normal" units by dividing
by $(\hbar c)^3$:

$$B = \frac{200^4 \text{ MeV}^4}{(197,327)^3 \text{ MeV}^3 \text{ fm}^3}$$
$$\approx \underline{\underline{208,24 \text{ MeV fm}^{-3}}}$$

Using the result from c, this gives
the size of the hadron as

$$R = \left(\frac{3}{16\pi} \frac{E_{H, \min}}{B} \right)^{1/3}$$
$$= \left(\frac{3}{16\pi} \frac{10^3 \text{ MeV}}{209,24 \text{ MeV} \cdot \text{fm}^{-3}} \right)^{1/3}$$

$$\approx \underline{\underline{0,66 \text{ fm}}}$$

e) The bag pressure can be found by noting that it acts like a cosmological constant: Its contribution to the energy density is $BV/V = B = \text{constant}$, independent of the volume.

$$\text{Therefore } P_B = -3Bc^2 = \underline{\underline{-B}}$$

Alternatively, from the 1st law of thermodynamics, with $dS = 0$

$$TdS = 0 = dE + PdV$$

$$\Rightarrow P = - \frac{dE}{dV},$$

o.e.,

$$P = - \left(\frac{\partial E}{\partial V} \right)_S = - \left(\frac{\partial (BV)}{\partial V} \right)_S = -B.$$

In the following, note that there was a mistake in the original version of the problem: The bag pressure should be included in the quark phase, note the hadronic phase.

Early on, in the quark-gluon plasma phase, we have 8 gluons with spin 1, so they are bosons. They are also massless, so each has 2 internal degrees of freedom. The u and d quarks are fermions, each with 2 internal degrees of freedom (spin $-\frac{1}{2}$), 3 degrees of freedom from colour (r, g, b), and each with their own antiparticle. Assuming equilibrium with the photons at temperature T, we therefore get

$$g_{*, QGP} = 8 \cdot 2 + \frac{7}{8} \cdot \underset{\substack{\uparrow \\ \text{u, d}}}{2} \cdot \underset{\substack{\uparrow \\ \text{r, g, b}}}{2} \cdot \underset{\substack{\uparrow \\ \text{spin}}}{2} \cdot \underset{\substack{\uparrow \\ \text{colour}}}{3}$$

$$= 37$$

For a relativistic gas, $P = \frac{1}{3} \rho c^2$, and

$$\rho c^2 = \frac{\pi^2}{30} g_* \frac{(k_B T)^4}{(\hbar c)^3}$$

Adding the (negative) bag pressure then gives

$$P_{QGP} = \frac{1}{3} \frac{\pi^2}{30} \cdot 37 \frac{(k_B T)^4}{(\hbar c)^3} - B$$

$$= \frac{37\pi^2}{90} \frac{(k_B T)^4}{(\hbar c)^3} - B$$

For the hadrons, the bookkeeping is easier. We have 3 of them, and they all have spin 0 (bosons), giving just one internal degree of freedom. Furthermore, π^+ is the antiparticle of π^- , and π^0 is its own antiparticle, so there is no additional contribution from antiparticles. Therefore

$$P_H = \frac{1}{3} \frac{\pi^2}{30} \cdot 3 \cdot \frac{(k_B T)^4}{(hc)^3}$$

$$= \frac{\pi^2}{30} \frac{(k_B T)^4}{(hc)^3}$$

f) At the phase transition we have

$$P_H(T_c) = P_{\text{QGP}}(T_c)$$

$$\Rightarrow \frac{\pi^2}{30} \frac{(k_B T_c)^4}{(hc)^3} = \frac{37\pi^2}{90} \frac{(k_B T_c)^4}{(hc)^3} - B$$

$$\Rightarrow \frac{\pi^2}{90} (37 - 3) \frac{(k_B T_c)^4}{(hc)^3} = B$$

$$\Rightarrow k_B T_c = \left(\frac{90}{34\pi^2} \right)^{1/4} [(hc)^3 B]^{1/4}$$

$$\approx \underline{\underline{140 \text{ MeV}}} \quad (144 \text{ MeV})$$

g) We use the relation between time and temperature derived in the lectures:

$$t = 2,423 g_*^{-1/2} \left(\frac{1 \text{ MeV}}{k_B T} \right)^2 s$$

in the radiation-dominated era. As we have seen, before the QCD transition, quarks and gluons contributed 37 relativistic degrees of freedom. In addition, at temperatures $k_B T \geq 140 \text{ MeV}$, photons are relativistic, and barely of the fermions, electrons, ~~muons~~ muons, and neutrinos are relativistic (and in equilibrium) up to this point.

This gives

$$g_* = 37 + \underset{\substack{\uparrow \\ \text{photons}}}{2} + \frac{7}{8} \left(\underset{\substack{\uparrow \\ e^\pm}}{2} \cdot \underset{\substack{\uparrow \\ \text{spin}}}{2} + \underset{\substack{\uparrow \\ \mu^\pm}}{2} \cdot \underset{\substack{\uparrow \\ \text{spin}}}{2} + \underset{\substack{\uparrow \\ \nu}}{3} \cdot \underset{\substack{\downarrow \\ \text{spin}}}{2} \cdot \underset{\substack{\downarrow \\ \text{spin}}}{1} \right)$$

$$= 39 + \frac{7}{8} \cdot 14 = 39 + \frac{49}{4} = \frac{205}{4}$$

$$\text{so } t = 2,423 \cdot \left(\frac{205}{4} \right)^{-1/2} \left(\frac{1 \text{ MeV}}{140 \text{ MeV}} \right)^2 s$$

$$\approx \underline{\underline{1,7 \cdot 10^{-5} \text{ s}}}$$

Problem 2

a) By definition

$$\eta = - \int_t^{\infty} \frac{cdt'}{a(t')}$$

$c/a(t')$ is always positive,
so $\int_{t_1}^{t_2} \frac{cdt'}{a(t')} > 0$ for any t_1, t_2

with $t_2 > t_1$. Therefore

$$\int_t^{\infty} \frac{cdt'}{a(t')} > 0,$$

so $\eta \leq 0$ and approaches 0

only in the limit $t \rightarrow \infty$.

We find $\frac{d\eta}{dt}$ from the fundamental theorem of calculus:

$$\begin{aligned} \frac{d\eta}{dt} &= \frac{d}{dt} \left[- \int_t^{\infty} \frac{cdt'}{a(t')} \right] = \frac{d}{dt} \left[\int_{\infty}^t \frac{cdt'}{a(t')} \right] \\ &= \frac{c}{a(t)} > 0 \end{aligned}$$

For $k=0$, the radial coordinate of the event horizon is

$$r_{EH}(t) = \int_t^{\infty} \frac{cdt'}{a(t')}$$

and the proper distance for it at time t is

$$\begin{aligned} d_{EH}(t) &= a(t) r_{EH}(t) = a(t) \int_t^{\infty} \frac{cdt'}{a(t')} \\ &= -a(t) \eta. \quad \text{q.e.d.} \end{aligned}$$

b) The area of the EH is

$$A_{EH} = 4\pi d_{EH}^2,$$

so

$$\frac{dA_{EH}}{dt} = \underbrace{8\pi d_{EH}}_{>0} \frac{d(d_{EH})}{dt},$$

and therefore $\frac{dA_{EH}}{dt} \geq 0 \Leftrightarrow \frac{d(d_{EH})}{dt} \geq 0$

Furthermore,

$$\frac{d(d_{EH})}{dt} = \frac{d(d_{EH})}{d\eta} \frac{d\eta}{dt}$$

and from a) we know that $\frac{d\eta}{dt} > 0$,

so $\frac{d(d_{EH})}{dt} \geq 0$ if and only if

$$\frac{d(d_{EH})}{d\eta} \geq 0, \text{ q. e. d.}$$

c)

FI with $k=0$:

$$H^2 = \frac{8\pi G}{3} \rho$$

Take the derivative with respect to t :

$$2H\dot{H} = \frac{8\pi G}{3} \dot{\rho}$$

The continuity equation gives

$$\dot{\rho} = -3H(\rho + \frac{P}{c^2}),$$

so

$$\begin{aligned} H\dot{H} &= \frac{4\pi G}{3} \left[-3H(\rho + \frac{P}{c^2}) \right] \\ &= -4\pi G H (\rho + \frac{P}{c^2}), \end{aligned}$$

and since $H > 0$,

$$\dot{H} = -4\pi G (\rho + \frac{P}{c^2}), \text{ q. e. d.}$$

$$d) \quad K \equiv \frac{1}{a} \frac{da}{d\eta} \equiv \frac{a'}{a}$$

Then

$$K' = \frac{a''a - a' \cdot a'}{a^2} = \frac{a''}{a} - \left(\frac{a'}{a}\right)^2$$

Also,

$$\dot{H} = \frac{d}{dt} \left(\frac{\dot{a}}{a} \right) = \frac{d}{d\eta} \left(\frac{\dot{a}}{a} \right) \frac{d\eta}{dt} \quad \underbrace{\frac{d\eta}{dt}}_{=c/a}$$

$$= \frac{c}{a} \frac{d}{d\eta} \left(\frac{1}{a} \frac{da}{d\eta} \frac{d\eta}{dt} \right) \quad \underbrace{\frac{d\eta}{dt}}_{=c/a}$$

$$= \frac{c}{a} \frac{d}{d\eta} \left(\frac{c}{a^2} \frac{da}{d\eta} \right)$$

$$= \frac{c}{a} \left(-\frac{2c da}{a^3 d\eta} \right) \frac{da}{d\eta} + \frac{c^2}{a^3} \frac{d^2 a}{d\eta^2}$$

$$= -\frac{2c^2}{a^3} \left(\frac{a'}{a} \right)^2 + \frac{c^2}{a^2} \frac{a''}{a}$$

$$= \frac{c^2}{a^2} \left[\underbrace{-\left(\frac{a'}{a}\right)^2}_{=-K^2} + \underbrace{\frac{a''}{a}}_{=K'} - \left(\frac{a'}{a}\right)^2 \right]$$

$$= \frac{c^2}{a^2} (K' - K^2)$$

From c) :

$$\frac{c^2}{a^2} (K' - K^2) = -4\pi G \left(\rho + \frac{p_r}{c^2} \right)$$

$$\Rightarrow K' - K^2 = -4\pi G \frac{a^2}{c^2} \left(\rho + \frac{p_r}{c^2} \right)$$

(Mistake in the problem text, but it doesn't matter)

By assumption $\rho + \frac{p_r}{c^2} \geq 0$,

and since $4\pi G \frac{a^2}{c^2} > 0$, this results

implies $K' - K^2 \leq 0$, q. e. d.

e) Since $d_{EM} = -a\eta$

$$\begin{aligned} \frac{d}{d\eta}(d_{EM}) &= -\frac{da}{d\eta}\eta - a \cdot \frac{d\eta}{d\eta} \\ &= -a'\eta - a \end{aligned}$$

So $\frac{d}{d\eta}(d_{EM}) \geq 0$

$$\Leftrightarrow -a'\eta - a \geq 0$$

$$\Leftrightarrow \underline{a'\eta + a \leq 0}$$

So $a'\eta \leq -a$

$$\Leftrightarrow \frac{a'}{a}\eta \leq -1,$$

and since $\eta \leq 0$,

$$\underline{\underline{K = \frac{a'}{a} \geq -\frac{1}{\eta}}}, \text{ q.e.d.}$$

f)

$$K^1 - K^2 \leq 0$$

$$\Rightarrow \frac{dK}{d\eta} \leq K^2$$

$$\Rightarrow \frac{dK}{K^2} \leq d\eta$$

$$\Rightarrow \int_K^{K_0} \frac{dK}{K^2} \leq \int_{\eta}^0 d\eta = -\eta$$

$$\Rightarrow \frac{1}{K} - \frac{1}{K_0} \leq -\eta$$

$$\Rightarrow -\frac{1}{K_0} + \frac{1}{K} \leq -\eta$$

$$\Rightarrow \underline{\underline{\frac{1}{K_0} - \frac{1}{K} \geq \eta}}, \text{ q.e.d.}$$

g) $K \geq -\frac{1}{r}$ follows from f) if we can
 Missing text in the problem

show that $K > 0$ and $K_0 = \infty$

Since $K = \frac{a'}{a} > 0$ follows from the fact that $\dot{a} > 0$ and $\frac{d\eta}{dt} > 0$, we only have to show that $K_0 = \infty$

Recall that

$$\gamma = - \int_t^\infty \frac{c dt'}{a(t')}$$

and look at $\lim_{a \rightarrow \infty, t \rightarrow \infty} dt' = \frac{da}{a}$

$$\begin{aligned} \int_t^\infty \frac{c dt'}{a(t')} &= \int_a^\infty \frac{c}{a} \frac{da}{a} = \int_a^\infty \frac{c}{a} \frac{da}{\frac{da}{dt} dt} \\ &= \int_a^\infty \frac{c}{a} \frac{da}{c a'} = \int_a^\infty \frac{da}{a'} \\ &= \int_a^\infty \frac{da}{a^2} < \infty, \end{aligned}$$

where the last inequality follows from the assumption that the event horizon exists.

Assume $\frac{1}{|K|}$ is bounded from below, so that $\frac{1}{|K|} > \varepsilon > 0$ for $a \rightarrow \infty, t \rightarrow \infty$ ($\eta \rightarrow 0$)

$$\text{Then } \int_a^\infty \frac{da}{a^2} > \varepsilon \int_a^\infty \frac{da}{a} = \varepsilon \left[\ln a \right]_a^\infty = \infty,$$

contradicting the existence of the event horizon.

So we must have $\varepsilon = 0$ and $K_0 = \infty$,

and $K \geq -\frac{1}{r}$ follows, which

proves that the area of the EH is non-decreasing.
 QED!