Solutions to exercises week 40 FYS2160

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Schroeder 4.1

a)

Net work done by gas during one cycle: $|W| = 2P_1V_1$ Heat absorbed (steps where $Q>0$): $Q_h = \frac{33}{2}$ $\frac{33}{2}P_1V_1$ Efficiency: $e = \frac{|W|}{Q}$ $\frac{|W|}{Q_h} = \frac{4}{33} = 12\%$

b)

Efficiency of heat engine: $e \leq 1 - \frac{T_c}{T_c}$ *T^h*

Use the ideal gas law, $PV = NkT$, to find the ratio of T_c to T_h . Lower left corner has the lowest temperature, $T = T_c$. Step A: *P* increases by a factor $2 \Rightarrow T$ increases by the same factor 2. Step B: *V* increases by a factor $3 \Rightarrow T$ increases by a factor 3. The highest temperature is then found in the upper right corner of the rectangular cycle in the PV-diagram, and is $T_h = 2 \cdot 3T_c = 6T_c$.

Maximal efficiency: $e_{\text{max}} = 1 - \frac{T_c}{6T}$ $\frac{T_c}{6T_c} = 5/6 = 83\%$

Schroeder 5.2

Look up $\Delta_f H$ and *S* for N₂, H₂ and NH₃ at the back of the book, calculate $\Delta G = \Delta H - T \Delta S$ for the reaction N₂ + 3H₂ → 2NH₃ and check with the tabulated value of $\Delta G(NH_3)$. The change in Gibbs free energy in this process is $2\Delta G(NH_3) = -32.90$ kJ.

Schroeder 5.23

a)

Writing $\Phi = U - TS - \mu N$ in terms of infinitesimal changes of the quantities involved:

$$
d\Phi = dU - TdS - SdT - \mu dN - Nd\mu.
$$

Inserting the thermodynamic identity for *dU* gives

$$
d\Phi = -PdV - SdT - Nd\mu.
$$

The partial derivatives of Φ wrt. *T*, *V* or μ is obtained by holding the two other quantities fixed:

$$
\left(\frac{\partial \Phi}{\partial T}\right)_{V,\mu} = -S, \qquad \left(\frac{\partial \Phi}{\partial V}\right)_{T,\mu} = -P, \qquad \left(\frac{\partial \Phi}{\partial \mu}\right)_{T,V} = -N.
$$

b)

As was done for *F* and *G* in chapter 5.2 in Schroeder, express an infinitesimal change in the total entropy S_{total} as a sum of changes in the entropy S of the system and the entropy S_R of the reservoir it is in contact with:

$$
dS_{\text{total}} = dS + dS_R.
$$

Assuming *V* fixed, apply the thermodynamic identity to dS_R (solve the expression for dU_R in terms of dS_R), and use the fact that $dU_R = -dU$ and $dN_R = -dN$ to write dS_R in terms of system variables. The total entropy of the system plus reservoir (expressed in terms of system variables) is then

$$
dS_{\text{total}} = -\frac{1}{T}(dU - TdS - \mu dN) = -\frac{1}{T}d\Phi.
$$

The tendency of the total entropy to increase therefore means that the grand free energy tends to decrease (since dS_{total} is positive, $d\Phi$ must be negative in order to make $-d\Phi/T$ positive).

c)

$$
\Phi = U - TS - \mu N
$$

= U - TS + PV - PV - \mu N
= G - PV - \mu N
= -PV,

where we in the last step use that $G = \mu N$ when *T* and *P* are constant (from the thermodynamic identity of *G*, $dG = -SdT + VdP + \mu dN$ with $dN = 1$.

d)

For the unoccupied state: $U = S = N = 0 \Rightarrow \Phi = 0$. For the occupied state: $U = U_0 = -13.6$ eV, $S = 0$, $N = 1$

$$
\Phi_{\text{occupied}} = U_0 - \mu
$$

= $U_0 + kT \ln \left[\frac{V}{N} \left(\frac{2\pi mkT}{h^2} \right)^{3/2} \right],$

where eq. 3.63 in Schroeder is used for the chemical potential of an ideal gas. Here *m* is the electron mass, and the electron concentration is given as $N/V = 2 \times 10^{19} \text{ m}^{-3}$. Putting in numbers gives us (the logarithm is calculated to be equal to 17.8)

$$
\Phi_{\text{occupied}} = -13.6 \text{ eV} + 8.62 \times 10^{-5} \text{ eV/K} \times 5800 \text{ K} \times 17.8
$$

$$
\approx -4.7 \text{ eV} < \Phi_{\text{unoccupied}} = 0 \text{ eV}.
$$

To find the temperature at which the occupied and unoccupied states are equally likely, we set their grand free energies to be equal and solve for T (neglecting the $\ln T^{3/2}$ dependence, which is insignificant compared to the factor T outside the logarithm, when T is varied):

$$
\Phi_{\text{occupied}} = \Phi_{\text{unoccupied}}
$$

$$
U_0 - 17.8kT = 0
$$

$$
T = -\frac{U_0}{17.8k} \approx 8800 \text{ K}
$$

Schroeder 5.32

a)

The Clausius-Clapeyron relation: $\frac{dP}{dT} = \frac{\Delta S}{\Delta V} = \frac{L}{T\Delta V}$ $\frac{L}{T\Delta V}$. When ice melts to water, the entropy increases $(\Delta S, L > 0)$, while the volume decreases $(\Delta V < 0)$; therefore, $dP/dT < 0$.

b)

Latent heat for converting ice into water: $L = 333 \text{ J/g}$ Density of ice: $\rho_{\rm ice} = 917 \text{ kg/m}^3 = 0.917 \times 10^6 \text{ g/m}^3$ Density of water: $\rho_{\text{water}} = 1000 \text{ kg/m}^3 = 1.000 \times 10^6 \text{ g/m}^3$ Volume of 1 g of ice: $V_{\text{ice}} = \frac{1 \text{ g}}{0.917 \times 10^6 \text{ g/m}^3} = 1.091 \times 10^{-6} \text{ m}^3$ Volume of 1 g of water: $V_{\text{water}} = 1.000 \times 10^{-6} \text{ m}^3$ $\Delta V = V_{\rm water} - V_{\rm ice} = -0.091 \times 10^{-6}~{\rm m}^3$

Ice melts at −1 ◦C if the pressure is at least

$$
\frac{dP}{dT} = \frac{L}{T\Delta V} = \frac{333 \text{ J}}{273 \text{ K} \times (-0.091 \times 10^{-6} \text{ m}^3)}
$$

$$
\approx -1.35 \times 10^7 \text{ Pa/K} = -135 \text{ bar/K}.
$$

c)

The pressure *P* (force per area) under an ice column of height *z* and density ρ is $P = \rho gz$. Therefore, the depth corresponding to a pressure of 135 bar is

$$
z = \frac{P}{\rho g} = \frac{135 \times 10^5 \text{ N/m}^2}{917 \text{ kg/m}^3 \times 9.81 \text{ N/m}^2} \approx 1500 \text{ m}.
$$

d)

As a crude approximation, take the area under the ice skate that's in contact with the ice to be ~ 1 mm in width and ~ 10 cm in length, that is $\sim 10^{-4}$ m². An ice skater weighing 50 kg would exert a force of ∼ 500 N on the ice. This amounts to a pressure on the ice of $P = 500 \text{ N}/10^{-4} \text{ m}^2 = 5 \times 10^6 \text{ Pa} = 50 \text{ bar}.$ Under this pressure the melting point only drops by \sim (50 bar)/(135 bar/°C) ≈ 0*.*4 ◦C. So the temperature needs to be within half a degree from the melting point in the first place, in order for this to have any effect.

Schroeder 6.5

a)

Consider a particle with three energy levels: $\epsilon_1 = -0.05$ eV, $\epsilon_2 = 0$ eV and $\epsilon_3 = 0.05$ eV. At 300 K, $kT \approx 0.026$ eV. The partition function is then

$$
Z = e^{-(-0.05/0.026)} + e^{0} + e^{-(0.05/0.026)} \approx 6.84 + 1 + 0.15 = 7.99.
$$

b)

The probability of an energy state ϵ_i is equal to

$$
P(\epsilon_i) = \frac{e^{-\epsilon_i/kT}}{Z},
$$

giving

$$
P(\epsilon_1 = -0.05 \text{ eV}) \approx 0.86
$$

$$
P(\epsilon_2 = 0 \text{ eV}) \approx 0.13
$$

$$
P(\epsilon_3 = 0.05 \text{ eV}) \approx 0.02
$$

c)

We now shift the zero point of energy so that: $\epsilon_1 = 0$ eV, $\epsilon_2 = 0.05$ eV and $\epsilon_3 = 0.10$ eV. In this case the partition function becomes

$$
Z = e^{0} + e^{-(0.05/0.026)} + e^{-(0.10/0.026)} \approx 1 + 0.15 + 0.02 = 1.17,
$$

with corresponding probabilities

$$
P(\epsilon_1 = 0 \text{ eV}) \approx 0.86
$$

$$
P(\epsilon_2 = 0.05 \text{ eV}) \approx 0.13
$$

$$
P(\epsilon_3 = 0.10 \text{ eV}) \approx 0.02
$$

The choice of zero-point of energy is arbitrary. Shifting the zero-point changes the values of the Boltzmann factors and the sum of these. However, the probability of the energy states $(P_i = e^{-\epsilon_i/kT}/Z$, that is, the ratio of individual Boltzmann factors to the sum) are not dependent on this choice.

Schroeder 6.6

Hydrogen atom in the ground state $(n = 1)$: $E_1 = -13.6$ eV.

Hydrogen atom in the first excited state $(n = 2)$: $E_2 = -3.4$ eV.

 $\Delta E = E_2 - E_1 = 10.2$ eV

Probability of finding a Hydrogen atom in one of its first excited states compared to the ground state:

$$
\frac{P(s_2)}{P(s_1)} = \frac{e^{-E(s_2)/kT}}{e^{-E(s_1)/kT}}
$$

Degeneracy of energy levels: $d(n) = \sum_{l=0}^{n-1} (2l + 1) = n^2$

There are four states with energy *E*2, hence finding a Hydrogen atom at room temperature (300 K) in a state with energy E_2 compared to the ground state becomes

$$
\frac{P(E_2)}{P(E_1)} = 4e^{-(E_2 - E_1)/kT} \approx 4e^{-10.2/0.026} \approx 10^{-171}.
$$

Hence, you will never find a Hydrogen atom in its first excited state at room temperature.

Now, doing the same calculation for a Hydrogen atom in the atmosphere of the star γ UMa, with surface temperature $T = 9500$ K:

$$
\frac{P(E_2)}{P(E_1)} = 4e^{-10.2 \text{ eV}/(k \times 9500 \text{ K})} \approx 1.6 \times 10^{-5}.
$$

In the atmosphere of γ UMa approximately 1 in 64 000 Hydrogen atoms will be found in one of their first excited states.

Compendium 8.1

a)

First law of thermodynamics: $\Delta U = Q + W$ Adiabatic process: $Q = 0 \Rightarrow \Delta U = Q \Rightarrow \Delta U - W = 0$ Finding expressions for ∆*U* and *W*:

Equipartition theorem: $U = \frac{1}{2}$ $\frac{1}{2}$ *fnRT* \Rightarrow $\Delta U = \frac{1}{2}$ 2 *fnR*∆*T*

Work: $W = -P\Delta V = -\frac{nRT}{V}\Delta V$ (the ideal gas law has been used in the last step)

Inserting the expressions for ∆*U* and *W*:

$$
\Delta U - W = 0
$$

$$
\frac{1}{2}fnR\Delta T - \left(-\frac{nRT}{V}\Delta V\right) = 0
$$

$$
\frac{fR}{2}\frac{\Delta T}{T} + R\frac{\Delta V}{V} = 0
$$

Now, what remains is to show that $c_{V,m} = fR/2$. In general the heat capacity can be expressed as

$$
C = \frac{Q}{\Delta T} = \frac{\Delta U - W}{\Delta T} = \frac{\Delta U + P\Delta V}{\Delta T}.
$$

Heat capacity at constant volume is then

$$
C_V = \left(\frac{\Delta U}{\Delta T}\right)_V = \frac{fnR\Delta T/2}{\Delta T} = \frac{fnR}{2},
$$

which for 1 mole of gas $(n = 1)$ becomes

$$
c_{V,m} = \frac{fR}{2}.
$$

Inserting this last relation, we end up with the expression we were asked to show:

$$
c_{V,m}\frac{\Delta T}{T} + R\frac{\Delta V}{V} = 0
$$

Let $\Delta T \to dT$ and $\Delta V \to dV$, and integrate the expression in a) over T and *V* between two states, 1 and 2:

$$
c_{V,m} \int_{T_1}^{T_2} \frac{dT}{T} = -R \int_{V_1}^{V_2} \frac{dV}{V}
$$

\n
$$
c_{V,m} \ln \frac{T_2}{T_1} = -R \ln \frac{V_2}{V_1} = R \ln \frac{V_1}{V_2}
$$

\n
$$
\ln \left(\frac{T_2}{T_1}\right)^{c_{V,m}} = \ln \left(\frac{V_1}{V_2}\right)^R
$$

\n
$$
\left(\frac{T_2}{T_1}\right)^{c_{V,m}} = \left(\frac{V_1}{V_2}\right)^R
$$
 (exponentiating both sides)
\n
$$
T_2^{c_{V,m}} V_2^R = T_1^{c_{V,m}} V_1^R
$$

\n
$$
T_2 V_2^{R/c_{V,m}} = T_1 V_1^{R/c_{V,m}}
$$
 (taking the $c_{V,m}$ -root of both sides)
\n
$$
T_2 V_2^{\alpha} = T_1 V_1^{\alpha},
$$

where $\alpha = R/c_{V,m}$.

c)

Using the ideal gas law to insert $T = \frac{PV}{nR}$:

$$
TV^{\alpha} = \frac{PV}{nR}V^{\alpha} = \frac{P}{nR}V^{\alpha+1}.
$$

b)

The heat capacity at constant pressure is

$$
C_P = \left(\frac{Q}{\Delta T}\right)_P = \left(\frac{\Delta U - W}{\Delta T}\right)_P
$$

=
$$
\left(\frac{fnR\Delta T/2 + P\Delta V}{\Delta T}\right)_P
$$

=
$$
\frac{fnR}{2} + P\left(\frac{\Delta V}{\Delta T}\right)_P
$$

=
$$
\frac{fnR}{2} + P\frac{\Delta (nRT/P)_P}{\Delta T}
$$

=
$$
\frac{fnR}{2} + P\frac{nR\Delta T/P}{\Delta T}
$$

=
$$
\frac{fnR}{2} + nR.
$$

The specific heat capacity for one mole at constant pressure, is then

$$
c_{P,m} = \frac{fR}{2} + R
$$

$$
= c_{V,m} + R
$$

$$
\Rightarrow \frac{c_{P,m}}{c_{V,m}} = 1 + \frac{R}{c_{V,m}} = 1 + \alpha \equiv \gamma
$$

This gives us

$$
TV^{\alpha} = \frac{P}{nR}V^{\alpha+1} = \frac{P}{nR}V^{\gamma}.
$$

For one mole

$$
T_1V_1^{\alpha} = T_2V_2^{\alpha}
$$

\n
$$
\frac{P_1}{R}V_1^{\gamma} = \frac{P_2}{R}V_2^{\gamma}
$$

\n
$$
P_1V_1^{\gamma} = P_2V_2^{\gamma}.
$$

Inserting $V = RT/P$ in the equation above, we can in a similar manner rewrite it on the following form

$$
T_1 P_1^{\beta} = T_2 P_2^{\beta},
$$

where $\beta \equiv (1 - \gamma)/\gamma$.

Compendium 8.2

a)

Isentropic $=$ adiabatic $+$ quasistatic. A quasistatic process means a change in volume that is sufficiently slow so that the gas has time to equilibrate to the changing conditions, and therefore maintains a uniform pressure. In order to compress the gas non-quasistatically, one would need to compress the gas faster than it is able to respond. This corresponds to a speed at least as fast as the speed of sound in the gas, and so quasistatic compression is often a reasonable assumption.

b)

For an ideal gas undergoing an adiabatic process, we know that $PV^{\gamma} = C$, where *C* is some constant. Solved for V, this becomes $V = (C/P)^{1/\gamma}$. The isentropic compressibility (with constant entropy *S*) is then

$$
\kappa_S = -\frac{1}{V} \left(\frac{\partial V}{\partial P} \right)_S = -\frac{1}{V} \frac{\partial}{\partial P} \left(\frac{C}{P} \right)^{1/\gamma}
$$

= $-\frac{C^{1/\gamma}}{V} \left(-\frac{1/\gamma}{P^{1/\gamma+1}} \right) = \frac{C^{1/\gamma}}{V} \left(\frac{1/\gamma}{P P^{1/\gamma}} \right)$
= $\frac{1}{\gamma PV} \left(\frac{C}{P} \right)^{1/\gamma} = \frac{1}{\gamma PV}$
= $\frac{1}{\gamma P}.$

For an isothermal process (*T* is constant) we use the ideal gas law to find an expression for *V* in terms of *P*: $V = (nRT)/P$. The isothermal compressibility is then

$$
\kappa_T = -\frac{1}{V} \left(\frac{\partial V}{\partial P} \right)_T = -\frac{1}{V} \frac{\partial}{\partial P} \left(\frac{nRT}{P} \right)
$$

= $-\frac{1}{V} \left(-\frac{nRT}{P^2} \right) = \frac{nRT}{P^2 V}$
= $\frac{1}{P^2} \frac{nRT}{V} = \frac{P}{P^2}$
= $\frac{1}{P}$.

The ratio

$$
\frac{\kappa_T}{\kappa_S} = \frac{1/P}{1/(\gamma P)} = \frac{\gamma P}{P} = \gamma,
$$

therefore tells us that κ_S is smaller than κ_T by a factor γ , as we wanted to show.

c)

The speed of sound in air can be expressed as $v = (\kappa_S \rho)^{-1/2}$, where $\kappa_S =$ $1/(\gamma P)$ is the isentropic compressibility and $\rho = nM/V$, with *n* equal to the number of moles, M the mass of one mole and V the total volume of the gas. Using this information we can write

$$
v = (\kappa_S \rho)^{-1/2} = \left(\frac{1}{\gamma P} \frac{nM}{V}\right)^{-1/2}
$$

$$
= \left(\frac{nM}{\gamma nRT}\right)^{-1/2} \qquad \text{(using } PV = nRT\text{)}
$$

$$
= \sqrt{\frac{\gamma RT}{M}}.
$$

d)

Using that $\gamma = (f + 2)/f$, where the number of active degrees of freedom f in air at room temperature $(T = 300 \text{ K})$ is assumed to be 5 (3 translational and 2 rotational), the speed of sound in air is calculated to be

$$
v = \sqrt{\frac{[(5+2)/5] \times 8.315 \text{ J/mol} \cdot \text{K} \times 300 \text{ K}}{2.9 \times 10^{-2} \text{ kg/mol}}} \approx 347 \text{ m/s}.
$$

e)

The average translational kinetic energy per molecule is

$$
\overline{K}_{\text{trans,molecule}} = \frac{1}{2} m \overline{v^2} = \frac{1}{2} m (\overline{v_x^2 + v_y^2 + v_z^2}) = \frac{3}{2} kT.
$$

For *n* moles of air, with molar mass *M*, this amounts to

$$
\overline{K}_{\text{trans}} = \frac{1}{2} n M \overline{v^2} = \frac{3}{2} n RT.
$$

Solving for v^2 gives

$$
\overline{v^2} = \frac{3RT}{M}.
$$

Taking the square root on both sides does not give us the average speed, but the root-mean-square (rms), which is a close approximation to the average:

$$
v_{\rm rms} \equiv \sqrt{\overline{v^2}} = \sqrt{\frac{3RT}{M}} \approx 508 \text{ m/s}.
$$