## FYS3500 - Problem set 2

### Spring term 2019

#### Problem 1 – In class

- a) What is the nuclear radius? Give several different ways to measure it and compare important differences.
- b) All known heavy nuclei exhibit an excess neutrons. Looking eg. at  $^{240}$ Pu, it has more then 40 % more neutrons then protons. Naively one should expect a larger radius for the neutrons then for the protons ( $\rightarrow$  How do we usually call these radii), however, they are observed to agree within about 0.1fm. How can this be explained?
- c) When analyze the scattering of  $\alpha$ -particles on medium and heavy nuclei we can learn about the nuclear radius. How?
- d) What is the angular dependence of scattered  $\alpha$ -particles? Sketch the cross-section of backscattered nuclei (scattering angle  $\theta \approx \pi$ ) as a function of the incident energy. Take this to estimate the nuclear radius of  $^{64}$ Cu, where the critical energy  $E_{\rm c}$  is about 13.7 MeV. Hint: What is the relation of kinetic and potential energy in backscattering?

#### Problem 2 – in class

When we describe nuclear  $\gamma$ -ray resonances, we usually give the energy  $E_{\gamma}$  of an emitted photon is the difference  $E_0$  between the excited state with energy  $E_x$  and the ground state (GS),  $E_{\gamma} = E_0 = E_x - E_{GS}$ . This is not exact for free atoms or molecules, as it neglects the recoil energy of the nucleus.

- a) Suppose that the nucleus was at rest before  $\gamma$ -ray emission and calculate the exact gamma-ray energy. Can we neglect the recoil effect?
- b)  $^{60}$ Co is one of the most important  $\gamma$ -ray calibration sources. By  $\beta$ -decay it feeds excited levels in  $^{60}$ Ni. The 1332 keV level (with direct decay into ground state) has a half-life  $t_{1/2}$  below 1ps ( $< 10^{-12}$  s). What does this imply for the energy of the  $\gamma$ -ray emitted in the emission? (Remember to convert half-life to lifetime)
- c) What are the implication for the nuclear resonance absorption of  $\gamma$ -ray photons? Argue qualitatively how these results would be effected if the nucleus was not at rest during decay.
- d) (At home?) What is the Mössbauer effect and how does this combine with these results?

# Problem 3: Nuclear binding energy

The emiprical mass formula is (neglecting the odd-even effect)

$$M(A,Z) = Z(m_p + m_e) + (A - Z)m_n - a_v A + a_s A^{2/3} + a_c \frac{Z(Z - 1)}{A^{1/3}} + a_{\text{sym}} \frac{(A - 2Z)^2}{A}$$
(1)

Where  $a_v$ ,  $a_s$ ,  $a_c$  and  $a_{sym}$  are constants. Explain the dependence on Z and A for the different terms

a) Two isobar nuclei 1 and 2 are called mirror nuclei if they result in each other by the exchange of the neutron an proton number:  $A_1 = A_2$ ,  $Z_1 = N_2$ ,  $N_1 = Z_2$ . They have a analogous structure and therefore also the same quantum numbers for the total angular momentrum J and parity  $\pi$ . As an example we look at <sup>27</sup>Al and <sup>27</sup>Si, both having ( $J=5/2,\pi=+1$ ) in the ground state. <sup>27</sup>Si decays by  $\beta^+$ -decay into <sup>27</sup>Al (<sup>27</sup>Si  $\rightarrow$ <sup>27</sup>Al+ $e^+$  +  $\nu$ ), where the sum of the kinetic

energies of the positron and neutrino has been measured to be at max  $E_0 = 3.8$  MeV. Deduce the radius parameter  $r_0$  from  $E_0$ .

Hint: If we assume that the nuclear forces are charge independent, the energy released in the decay of mirror nuclei is given by the mass difference  $(m_n - m_p)^2 = 1.29$  MeV and the difference in the coulomb energy. The charge distribution can be assumes as a homogeneously charged sphere (total charge Ze and radius R) with the coulomb energy  $W_c$ 

$$W_c = \frac{3}{5} \frac{Q^2}{4\pi\epsilon_0} \frac{1}{R} \tag{2}$$

b) Use (1) to show that for a constant *A* the *Z* value that corresponds to the most stable nucleus is

$$Z_{\min} = \frac{m_n c^2 - (m_p + m_e)c^2 + a_c A^{-1/3} + 4a_{\text{sym}}}{2a_c A^{-1/3} + 8a_{\text{sym}} A^{-1}}$$

c) Determine the most stable isobar with mass number A=87. (Again, use  $a_v=15.5$  MeV,  $a_s=16.8$  MeV,  $a_c=0.72$  MeV and  $a_{\rm sym}=23$  MeV).

## **Problem 4: Reaction Q-values**

Find the Q-value of the reaction:

$$^{86}\text{Kr} + \text{d} \rightarrow ^{87}\text{Kr} + \text{p}$$

Rember:  $d = {}^{2}H$ .

# Problem 5: "The binding energy plot"

Explain how the mass of a nucleus can be calculated from the plot in figure below. Explain briefly some of the main features of the plot and estimate the mass of <sup>130</sup>Xe. What is the relation to fission and fusion? How did we get the heavy elements?

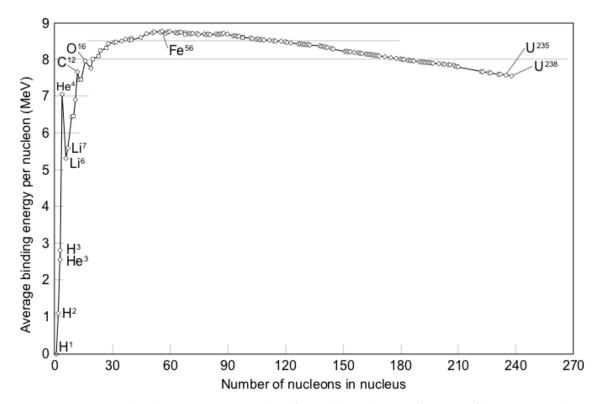


Figure 1: Average binding energy per nucleon for stable nuclei as a function of the mass number.

## **Problem 6: Nuclear Scattering**

In Figure 2 the differential cross-sections  $d\sigma/d\Omega$  for scattering of high energy electrons on  $^{40}$ Ca and  $^{48}$ Ca are displayed.

- a) Explain the general behavior of the cross-section as a function of the scattering angle. What is the source for the (local)minima?
- b) The form factor  $F(q^2)$  is defined as the Fourier transformation of the charge density  $\rho(r)$ , with the momentum transfer  $q=k_{\rm f}-k_{\rm i}$  and the initial and final momentum  $k_{\rm i,f}$ . Under the usual assumption of the Born approximation and negligible recoil,  $F(q^2)$  can be calculated by

$$F(q^2) = \int e^{iq\mathbf{r}/\hbar} \rho(\mathbf{r}) d^3 \mathbf{r}.$$
 (3)

Calculate  $F(q^2)$  for a *homogeneous* spherical charge distribution,  $\rho(r) = 3Ze/(4\pi R^3)$  for r < R, and 0 elsewhere. Show that the result is given by

$$F(q^2) = 3Ze(\frac{\hbar}{qR})^3 \left[ \left( \frac{qR}{\hbar} \right) \cos(\frac{qR}{\hbar}) - \sin(\frac{qR}{\hbar}) \right] \tag{4}$$

- c) Plot the result. (It is important to choose a reasonable scaling for q. Recall that  $|q| = 2|p|\sin(\frac{\theta}{2})$ )
- d) Use this and Figure 2 to compare the radius of  $^{40}$ Ca and  $^{48}$ Ca. Hint: How does the location of the minima depend on R for given angle.

e) Some extra calculations ("bonus"): Develop eq.(3) for a general spherical potential in powers of |q| (first 2 non-vanishing terms) using the mean square radius  $\langle r^2 \rangle$ . This will lead you to

$$F(q) = 1 - \frac{1}{6} \frac{q \langle r^2 \rangle}{\hbar^2} + \dots$$
 (5)

How can you solve the expression for the mean square charge radius  $\langle r^2 \rangle$ . What momentum transfers are most important in order to measure  $\langle r^2 \rangle$ ? What does this mean for the experiment (for example, would you measure at certain angles, rather high or low energies,...?) Hint: a) The mean square charge radius is defined as  $\langle r^2 \rangle = \int r^2 \rho(r) \, d^3 r$ . b) Remember that the form factor is normed such that  $F(q^2 = 0) = 0$ .

f) For further thought ;P: Experimentally only a restricted range of momentum transfers is accessible, as it is limited by the beam intensity. In addition, the cross-section drops quickly with increasing momentum transfers. Think about a method to determine the charge distribution despite these problems.

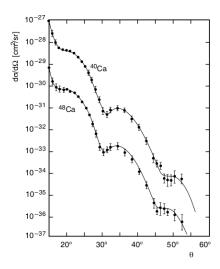


Figure 2: Differential cross-section for scattering of 750 MeV electrons on  $^{40}$ Ca and  $^{48}$ Ca. The cross-sections have been multiplied by 10 for  $^{40}$ Ca and by  $10^{-1}$  for  $^{48}$ Ca for displaying purposes. Source: J. B. Bellicard, et al. Phys. Rev. Lett. **19** (1967) 527

# Problem 7: LS-coupling in the shell model – Special hand-in for this part: In 2 weeks

In order to get the correct magic numbers in the nuclear shell model we need to include a spinorbit coupling. This leads to a Hamiltonian similar to:

$$\hat{H} = \hat{H}_0 - V_{SO}\hat{\vec{L}} \cdot \hat{\vec{S}} \tag{6}$$

Where  $\hat{H}_0$  is a Hamiltonian with eigenstates  $\hat{H}_0 | N, l \rangle = \hbar \omega (N+3/2) | N, l \rangle$  with  $l = N, N-2, N-4, \ldots 1$  or  $0, l \geq 0$  and  $V_{SO}\hat{\vec{L}} \cdot \hat{\vec{S}}$  is the spin-orbit coupling.  $V_{SO}$  is the strength of the coupling and can be regarded as a constant in this problem. In this problem we only look at spin-1/2 fermions.

a) The operator for the total spin is  $\hat{\vec{J}} = \hat{\vec{L}} + \hat{\vec{S}}$ . Where  $\hat{\vec{L}}$  is the angular momentum operator and  $\hat{\vec{S}}$  is the spin operator. Show that:

$$\hat{\vec{L}} \cdot \hat{\vec{S}} = \frac{1}{2} (\hat{\vec{J}}^2 - \hat{\vec{L}}^2 - \hat{\vec{S}}^2)$$

b) For a given angular momentum l the total spin state  $|j,m_j\rangle$  can be j=|l-1/2|,l+1/2. The two possible states can be written as a superposition of  $|l,m_l=m_j\pm 1/2\rangle\otimes |s=1/2,m_s\mp 1/2\rangle$  states:

$$\begin{split} |j = l \pm 1/2, m_j \rangle &= C_{1/2}^l |l, m_j - 1/2 \rangle \otimes |\uparrow \rangle \\ &+ C_{-1/2}^l |l, m_j + 1/2 \rangle \otimes |\downarrow \rangle \,, \end{split}$$

where the Clebsch-Gordan coefficients  $C_{m_s}^l=0$  if  $|m_j-m_s|>l$  or  $m_j+m_s>l$ ; and  $|\uparrow\rangle=|1/2,1/2\rangle$ ,  $|\downarrow\rangle=|1/2,-1/2\rangle$ . Find the energy of the state  $|N,l,j=l\pm 1/2,m_j\rangle$  using (6).