

a) we find

$$\psi = R(r) Y_{l,m_l}(\theta, \varphi)$$

which is given by

$\bullet R_{nl}(r) = N_{nl} r^l e^{-\nu r^2} L_n^{(l+\frac{1}{2})} (2\nu r^2)$

$\uparrow$   
some normalization factor       $\uparrow$   
"Legendre Polynomials"

$\circ$  or equivalently  
~~through~~ through Hermitian Polynomials

$\bullet Y_{l,m_l} \rightarrow$  Spherical harmonics

$$= \sqrt{(2l+1) \frac{(l-m)!}{(l+m)!}} P_l^m(\cos\theta) e^{im\varphi}$$

$\underbrace{\hspace{10em}}$   
Legendre polynomials

b) possible values for  $l$  are

$$l = N, N-2, \dots, \begin{cases} 1 \\ 0 \end{cases}, \text{ so for } \begin{array}{l} N=5 \rightarrow l = \{1, 3, 5\} \\ N=6 \rightarrow l = \{0, 2, 4, 6\} \end{array}$$

c)

N	$E_N (\text{keV})$	(n, l)	spin $\frac{1}{2}$ and $\frac{3}{2}$		total
			$\sum_{(n,l)} 2(2l+1)$	total	
0	$3/2$	$1s$	2	2	
1	$5/2$	$2p$	6	8	
2	$7/2$	$2s, 1d$	12	20	
3	$9/2$	$2p, 1f$ <del>3s, 2d, 1g</del>	20	40	
4	$11/2$	$3s, 2d, 1g$	30	70	

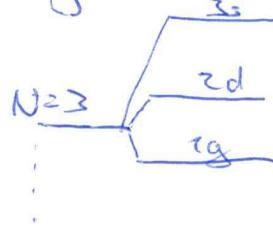
$N$	$\langle L^2 \rangle_N$
0	0
1	$\frac{1+4}{2} = 2$
2	$\frac{2+5}{2} = 5$
3	$\frac{3+6}{2} = 9$

$N$	$L \xrightarrow{\text{LLtot}}$	$L^2 \xrightarrow{} \langle L^2 \rangle_N$
0	$1s \rightarrow 0$	$0-0=0$
1	$2p \rightarrow \frac{1(1+1)}{2} = 1$	$2-2=0$
2	$2s \rightarrow 0$ $1d \rightarrow \frac{2(2+1)}{6} = 1$	$0-5=-5$ $6-5=1$
3	$3s \rightarrow 0$ $2d \rightarrow 6$ $1g \rightarrow \frac{2(2+1)}{12} = 1$	$0-9=-9$ $6-9=-3$ $12-9=3$

$\Rightarrow$  splitting of  
2s and 1d!

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dito

$\Rightarrow$   $a_1$  must be negative in order to have higher  $l$  values at lower (more bound) energies



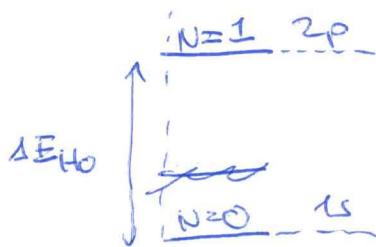
(however, note that the usual convention is to chose)

$$H_{\text{HO}} = -\mu (L^2 + L^2_{\text{tot}}) \text{thw, where}$$

or

$a_1,$

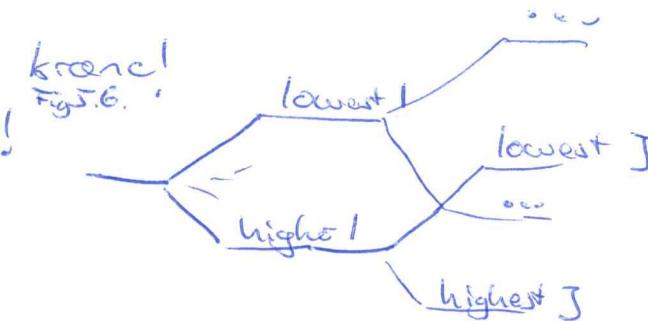
$\text{so } \mu > 0$



the relation of the splittings of the HO and HHO addition ( $L^2_{\text{tot}}$ ) is given by the ratio of the coefficients / ~ "how large is  $\mu$ "?!

a)

~ see  
important!



[2]  
+  
deg. in HO  
B

1s<sub>1/2</sub>  
Spectroscopic  
notation

[2] 2  
deg. in this  
level

total number of  
nucleons + 1 here  
(protons or neutrons)

• note: degeneracy here for protons ~~or~~ neutrons!  
= 2 × for both of them

b) Magic numbers: 8, 20, 28, 50, 82, 126

↳ all of them predicted  $\rightarrow$

c)  $^{27}\text{Al}$  odd proton, hole in  $(1d_{5/2})^1 \rightarrow I^\pi = \frac{5}{2}^+$ ; ~~for~~

$^{28}\text{Si} \rightarrow I^\pi = 0^+$

$^{29}\text{Si} \rightarrow$  1 neutron in  $2s_{1/2} \rightarrow \frac{1}{2}^+$

$^{31}\text{P} \rightarrow$  1 p in  $2s_{1/2} \rightarrow \frac{1}{2}^+$

$^{32}\text{Si} \rightarrow 0^+$

$^{33}\text{Si} \rightarrow$  1 n in  $d_{3/2} \rightarrow \frac{3}{2}^+$

$^{41}\text{Ar} \rightarrow$  1 p in  $2d_{5/2} \rightarrow \frac{5}{2}^+$

same for e.g.  $^{45}\text{Ti}_{23} \rightarrow (2p, 3n)$  beyond

$Z=20$  but we need  
 $N=20$ ! to care only about  
the unpaired neutron  
 $\rightarrow$  in  $1f_{5/2} \Rightarrow I^\pi = \frac{7}{2}^-$

~~$^{26}\text{Al} \rightarrow$  look only at the unpaired neutron  $(1d_{5/2})^3$~~

d)  $^{23}_{11}\text{Na}$  - expected spin (naively)

core only about ~~less~~ unpaired ~~proton~~,  $\rightarrow \frac{1}{2}d\frac{5}{2} \rightarrow I^{\pi} = \frac{5}{2}^+$

however:  $I_{\text{exp}}^{\pi} = \frac{3}{2}^{\pm}$  (+?) — check up online

incomplete subshell, residual forces (beyond extremely, indep. single part! shell modell)

$^{203}_{81}\text{Ti}$  - again, one would just regard the hole in  $h_{11/2} \rightarrow I^{\pi} = \frac{11}{2}^-$

however:  $I_{\text{exp}}^{\pi} = \frac{1}{2}^+$   $\Rightarrow$  due to pairing:

pairing energy increases with l, such that it is more favourable to have this config. [last three shells given]

(actual)  $(2d\frac{3}{2})^4 (3s\frac{1}{2})^2 (1h\frac{11}{2})^1$  — particle number in that shell

then

even though h has a higher energy than s.

(<sub>ext. int.</sub> <sub>single part</sub>)  $(2d\frac{3}{2})^4 (3s\frac{1}{2})^2 (1h\frac{11}{2})^{\frac{11}{2}}$

→ extremely independent single particle shell modell is just one extreme and cannot explain all effects. We need to take into account e.g.

• deformation      • Many-particle shell model!

~~attraction~~

in more realistic models.

However, it works incredibly well for a large number of nuclei, including ~~most~~ <sup>most</sup> all ~~of~~ even-even, even-uneven nuclei in

$A < 150$  &  $190 < A < 220$

[ $\rightsquigarrow$  otherwise often large deformation]

$$\begin{aligned} &= (n_r - 1) + l \\ &\quad (0 = n_r' + l) \end{aligned}$$

$$V_{\text{NLS}} = \frac{M\omega_0^2}{2} r^2 - K \frac{\hbar}{m} \omega_0 [2\vec{l} \cdot \vec{s} + \mu(l^2 - \langle l^2 \rangle_N)]$$

e) all terms

$$\Rightarrow E_{N,j} = \hbar\omega_0 \left[ N + \frac{3}{2} + K \left\{ \begin{array}{l} l \\ -l-1 \end{array} \right\} - \mu \left( l(l+1) - \frac{N(N+3)}{2} \right) \right] \text{ for } \begin{cases} j = l + \frac{1}{2} \\ j = l - \frac{1}{2} \end{cases}$$