# Lecture 7

#### 06.02.2018

Classical free particles Maxwell-Boltzmann distribution

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#### Module II: Non-interacting particles, multiplicity function, partition function

on. 6. feb.	Classical free particles, Maxwell-Boltzmann distribution
fr. 8. feb.	Quantum ideal gases, Bose-Einstein distribution
on. 13. feb.	Fermi-Dirac distribution
fr. 15. feb.	Summary and questions

#### Free particles

• Mutual interactions between particles is negligible:

- > ideal spin systems (*paramagnetism*) -- **distinguishable particles**
- ideal classical gases
- ideal quantum gases

— > indistinguishable particles

#### Free spins

- Consider a paramagnetic solid composed of N <u>identical</u> spin ½ particles <u>localized</u> on a lattice
- Each particle is described by a spin  $s = \pm \frac{1}{2}$  and a magnetic moment  $m = 2\mu s = \pm \mu$ , where  $\mu$  is the Bohr magneton.



#### Free spins: microstates

• Each particle is described by a spin  $s = \pm \frac{1}{2}$  and a magnetic moment

 $m = 2\mu s = \pm \mu$ , where  $\mu$  is the Bohr magneton.

• Spin in a magnetic field *B* has a potential energy

$$\epsilon = -\boldsymbol{m} \cdot \boldsymbol{B} = -2\mu sB$$



#### Free spins: equilibrium macrostate

• Number of spins N

$$N = n_+ + n_-$$

• Total internal energy:

$$E = -\mu B(n_+ - n_-)$$

• Total magnetic moment:

$$M = \mu(n_+ - n_-)$$



#### Microcanonical ensemble of free spins

$$N = n_{+} + n_{-},$$
  $U = -\mu B(n_{+} - n_{-}),$   $M = \mu(n_{+} - n_{-})$ 

Free spins+ the external magnetic field = isolated system

Total energy is conserved. The <u>equilibrium macrostate</u> is determined by total (discrete) number of possible spin configurations (microstates) or its **multiplicity** W(U, N) (the equivalant of the phase space volume for systems with a continuum number of microstates).



### Basic rules of combinatorics

• Number of ways (permutations) in which we can arrange N *distrinct* objects in a *row* 

$$N \cdot (N-1) \cdots 2 \cdot 1 \equiv N!$$

• Ordered selection of r out of N distinct objects

$$N \cdot (N-1) \cdots (N-r+1) \equiv \frac{N!}{(N-r)!}$$

• When the selection is **unordered**,  $\frac{N!}{r!(N-r)!}$ 

# Microcanonical ensemble of free spins

Number of spin configurations with  $n_+$  spins up and  $n_-$  spins down out of N spins is

$$W(n_{+}, N) = \frac{N!}{n_{+}! n_{-}!} = \frac{N!}{n_{+}! (N - n_{+})!}$$

#### Microcanonical ensemble of free spins: Macrostate

• Boltzmann' entropy  $S = k \ln W$ 

$$S(n_+, N) = k(N \log N - n_+ \ln n_+ - (N - n_+) \ln(N - n_+))$$

• Energy 
$$U(n_+, N) = -\mu B(2n_+ - N)$$

• Temperature

$$= -\mu B (2n_{+} - N)$$

$$\frac{1}{T} = \left(\frac{\partial S}{\partial U}\right)_{N} = \frac{\partial S}{\partial n_{+}} \frac{\partial n_{+}}{\partial U} = k \ln \left(\frac{n_{+}}{N - n_{+}}\right) \frac{1}{2\mu B}$$

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$$\frac{n_+}{N-n_+} = e^{2\beta\mu B}$$

# Microcanonical ensemble: Maxwell-Boltzmann distribution

$$\frac{n_{+}}{N} = \frac{e^{\frac{\mu B}{kT}}}{e^{\frac{\mu B}{kT}} + e^{-\frac{\mu B}{kT}}}, \qquad \frac{n_{-}}{N} = \frac{e^{-\frac{\mu B}{kT}}}{e^{\frac{\mu B}{kT}} + e^{-\frac{\mu B}{kT}}}$$

- As temperature T gets lower, the distribution of spins up increases
- Thermal fluctuations increase the likelyhood that spins may align in the opposite directions with the applied magnetic field



#### Thermodynamics of paramagnets

$$\frac{n_{+}}{N} = \frac{e^{\frac{\mu B}{kT}}}{e^{\frac{\mu B}{kT}} + e^{-\frac{\mu B}{kT}}}, \qquad \frac{n_{-}}{N} = \frac{e^{-\frac{\mu B}{kT}}}{e^{\frac{\mu B}{kT}} + e^{-\frac{\mu B}{kT}}}$$

• Total magnetization

$$M = \mu(n_+ - n_-) = N\mu \tanh \frac{\mu B}{kT}$$

• Susceptibility

$$\chi = \left(\frac{\partial M}{\partial B}\right)_T = \frac{N\mu^2}{kT} \frac{1}{\cosh^2\left(\frac{\mu B}{kT}\right)}$$



#### Canonical ensemble of free spins at fixed T and N

Energy change upon flipping a +1/2 spin into a -1/2 spin due to thermal fluctuations:  $n_+ \rightarrow n_+ - 1$ ,  $n_- \rightarrow n_- + 1$ ,

$$U_1 = -\mu B(n_+ - n_-) \to U_2 = -\mu B(n_+ - 1 - n_- - 1)$$

 $\Delta U = 2\mu B$ 

Change in entropy upon a spin flipping

$$\Delta S = k(\log W_2 - \log W_1) = k\left(\log \frac{N!}{(n_+ - 1)! (n_- + 1)!} - \log \frac{N!}{(n_+)! (n_-)!}\right)$$
$$\Delta S \approx k \log \frac{n_+}{n_-}$$

Changes in energy and entropy are linked by the fixed temperature of the thermal bath

$$\frac{\Delta S}{\Delta U} = \frac{1}{T} \rightarrow \frac{n_{+}}{n_{-}} = e^{\frac{2\mu B}{kT}} \rightarrow \frac{n_{+}}{N} = \frac{e^{\frac{\mu B}{kT}}}{e^{\frac{\mu B}{kT}} + e^{-\frac{\mu B}{kT}}}$$

### Canonical ensemble of free spins: MBD

• Maxwell-Boltzmann distribution (MBD): probability that a spin occupies a microstate//fraction of spins in a (single-particle) microstate

$$\frac{n_s}{N} = \frac{1}{Z_1(T)} e^{-\frac{\epsilon_s}{kT}}, \qquad \epsilon_s = -s\mu B, \qquad s = \pm 1$$

• 1-particle partition function  $Z_1(T) = \sum_{s=\pm 1} e^{-\frac{\epsilon_s}{kT}} = 2 \cosh\left(\frac{\mu B}{kT}\right)$ 

• N-partition function 
$$Z_N(T) = Z_1^N = 2^N \cosh^N\left(\frac{\mu B}{kT}\right)$$

#### Canonical ensemble of free spins: macrostate

• Average spin energy 
$$\langle \epsilon \rangle = \frac{1}{Z_1} \sum_{s} \epsilon_s e^{-\frac{\epsilon_s}{kT}} = -\mu B \tanh\left(\frac{\mu B}{kT}\right)$$

- Average total energy  $U = N\langle \epsilon \rangle = -N\mu B \tanh\left(\frac{\mu B}{kT}\right)$
- Helmholtz free energy  $F(T, M(B, T)) = -NkT \log \left( 2 \cosh \left( \frac{\mu B}{kT} \right) \right)$
- Gibbs free energy G(T, B) = F MB

#### Maxwell-Boltzmann statistics: generic case

• Consider a generic ensemble of N free and distinguishable particles that can occupy discrete energy levels  $\{\varepsilon_i\}$  with the occupation numbers  $\{n_i\}$ 

$$N = \sum_{i} n_{i}, \qquad U = \sum_{i} n_{i} \epsilon_{i}$$

• Each energy state  $\epsilon_i$  has a «degeneracy»  $g_i$  which is the density of states, i.e. density of microstates at that energy level

## Microcanonical: Multiplicity of a macrostate

• Number of configurations with  $n_1$  particles in the energy state  $\epsilon_1$ ,  $n_2$  particles in the energy state  $\epsilon_2$ ,  $\cdots$  etc

Each particle in ε<sub>i</sub> energy level has g<sub>i</sub> available microstates, hence g<sub>i</sub><sup>n<sub>i</sub></sup> ways of arranging n<sub>i</sub> participles in g<sub>i</sub> degenerate states.

 $N! \prod_{i} \frac{1}{n_i!}$ 

• The total number of microstates with the particle configurations  $\{n_i\}$ 

$$W(\{n_i^{(eq)}\}) = N! \prod_i \frac{g_i^{n_i}}{n_i!}$$

• Total multiplicity of a macrostate is dominated by the equilibrium distribution, hence

$$\mathbf{S} = k \log \left( W(\{n_i^{(eq)}\}) \right)$$

#### Canonical ensemble

• Thermal fluctuation: a particle changes its energy from  $\epsilon_i$  to  $\epsilon_i$ 

$$n_i \rightarrow n_i + 1$$
,  $n_j \rightarrow n_j - 1$ 

• Change in energy  $\Delta U = \epsilon_i - \epsilon_j$ 

• Change in entropy 
$$\Delta S = k \left( \log \frac{g_i^{n_i+1}}{(n_i+1)!} \frac{g_j^{n_j-1}}{(n_j-1)!} - \log \frac{g_i^{n_i}}{n_i!} \frac{g_j^{n_j}}{n_j!} \right) \approx k \log \frac{g_i n_j}{n_i g_j}$$

$$\frac{\Delta S}{\Delta U} = \frac{1}{\mathrm{T}} \to \frac{g_i n_j}{n_i g_j} = \mathrm{e}^{-\beta(\epsilon_i - \epsilon_j)}$$

#### Maxwell-Boltzmann distribution

$$n_i = \frac{N}{Z_1} g_i \mathrm{e}^{-\beta \epsilon_i}, \ Z_1 = \sum_i g_i e^{-\beta \epsilon_i}$$

• Probability to find a particle in the energy level  $\epsilon_i$ :  $p_i = \frac{g_i}{Z_1} e^{-\beta \epsilon_i}$ 

• Average energy

$$\mathbf{U} = \mathbf{N} \langle \epsilon_i \rangle = N \sum_i p_i \epsilon_i = \frac{N}{Z_1} \sum_i \epsilon_i g_i e^{-\beta \epsilon_i} = -\frac{\partial}{\partial \beta} \ln Z_1^N$$

• Helmholtz free energy

$$F = -kT \ln Z_1^N$$

• Entropy

$$\frac{S}{k} = N \log Z_1 + \frac{U}{kT}$$

## Indistinguishable, free particles particles

• Number of configurations  $W_{\text{disting}} = N! \prod_{i} \frac{g_{i}^{n_{i}}}{n_{i}!}$ 

$$W_{\text{indisting}} = \prod_{i} \frac{g_{i}^{n_{i}}}{n_{i}!}$$

# Ideal gas

• Find the energy levels  $\epsilon_s$  for free particles in a box, so we can compute the partition function using the generic formula from M-B statistics

$$Z_N = \frac{Z_1^N}{N!}, \qquad \qquad Z_1 = \sum_{S} e^{-\beta \epsilon_S}$$

- Hamiltonian of one particle  $\hat{H} = \frac{\hat{p}^2}{2m}$  with the momentum operator  $\hat{p} = -i\hbar\nabla$
- The particle's wavefunction satisfies the *Schrödinger* equation

$$\widehat{H}\psi = \epsilon\psi \to -\frac{\hbar^2}{2m}\nabla^2\psi = \epsilon\psi$$



Ideal gas

#### Eigenfunctions of the *Schrödinger* equation

 $-\frac{\hbar^2}{2m}\nabla^2\psi=\epsilon\psi$  with periodic boundary conditions

• 
$$\psi_{n} = e^{2\pi i \frac{n \cdot x}{L}}$$
,  $\boldsymbol{n} = (n_{x}, n_{y}, n_{z})$ ,  $n_{i} = 0, \pm 1, \pm 2, \cdots$ 

• Energy levels 
$$\epsilon_n = \frac{\hbar^2}{2m} \left(\frac{2\pi}{L}\right)^2 \left(n_x^2 + n_y^2 + n_z^2\right)$$



## Ideal gas: $Z_1$

Energy levels  $\epsilon_n = \frac{\hbar^2}{2m} \left(\frac{2\pi}{L}\right)^2 \left(n_x^2 + n_y^2 + n_z^2\right)$ 

$$Z_{1} = \sum_{n_{\chi}} \sum_{n_{y}} \sum_{n_{z}} e^{-\beta \frac{\hbar^{2}}{2m} \left(\frac{2\pi}{L}\right)^{2} \left(n_{\chi}^{2} + n_{y}^{2} + n_{z}^{2}\right)} = \left(\sum_{n} e^{-\beta \frac{\hbar^{2}}{2m} \left(\frac{2\pi}{L}\right)^{2} n^{2}}\right)^{3}$$

$$Z_1 \approx \left( \int_{-\infty}^{\infty} dn \, e^{-\beta \frac{\hbar^2}{2m} \left(\frac{2\pi}{L}\right)^2 n^2} \right)^3 = L^3 \left(\frac{m}{2\pi\beta\hbar^2}\right)^{\frac{3}{2}}$$

$$Z_1 = \frac{V}{\Lambda^3}$$
, thermal wavelength  $\Lambda = \sqrt{\frac{2\pi\beta\hbar^2}{m}}$ 

#### Ideal gas

$$Z_1(T) = \frac{V}{\Lambda^3(T)}$$
, thermal wavelength  $\Lambda = \sqrt{\frac{2\pi\beta\hbar^2}{m}}$ 

• Maxwell-Boltzmann statistics is valid in the classical limit:

$$\Lambda^{3}(T) \ll \frac{L^{3}}{N} \to T \gg \left(\frac{h^{2}}{2\pi mk}\right) \frac{N^{\frac{2}{3}}}{L^{2}}$$

#### Ideal gas: (p,q)-phase space vs. n-space

Connection to the integral over the momentum phase space

$$\mathbf{Z}_1 = \int d^3 \boldsymbol{n} \, e^{-\beta \frac{\hbar^2}{2m} \left(\frac{2\pi}{L}\right)^2 \boldsymbol{n}^2}$$

The momentum coordinate  $p = \frac{2\pi\hbar}{L} n \rightarrow d^3 p = \frac{(2\pi\hbar)^3}{V} d^3 n$ 

$$Z_{1} = V \int \frac{d^{3} \boldsymbol{p}}{(2\pi\hbar)^{3}} e^{-\beta \frac{\boldsymbol{p}^{2}}{2m}} = 4\pi V \int \frac{p^{2} dp}{(2\pi\hbar)^{3}} e^{-\beta \frac{p^{2}}{2m}}$$

$$Z_{1} = 4\pi V \int \frac{p^{2} dp}{(2\pi\hbar)^{3}} e^{-\beta \frac{p^{2}}{2m}} = \frac{V}{\epsilon = \frac{p^{2}}{2m}} \frac{V}{4\pi^{2}} \left(\frac{2m}{\hbar^{2}}\right)^{\frac{3}{2}} \int d\epsilon \,\epsilon^{1/2} e^{-\beta\epsilon} = \int d\epsilon \mathrm{D}(\epsilon) e^{-\beta\epsilon}$$

Ideal gas: density of states

$$Z_1 = \int_0^\infty d\epsilon D(\epsilon) e^{-\beta\epsilon}$$
 (Laplace transform)

$$D(\epsilon) = \frac{V}{4\pi^2} \left(\frac{2m}{\hbar^2}\right)^{\frac{3}{2}} \epsilon^{1/2}$$

 $D(\epsilon)d\epsilon$  number of microstates with energy between  $\epsilon$  and  $\epsilon + d\epsilon$  (for 1 particle)

## Thermodynamics of the ideal gas

$$Z_{N} = \frac{Z_{1}^{N}}{N!} = \frac{V^{N}}{N! \Lambda^{3N}}$$

• Helmholtz free energy

• Pressure 
$$P = -\left(\frac{\partial F}{\partial V}\right)_{T,N} = \frac{NkT}{V}$$

• Chemical potential 
$$\mu = -\left(\frac{\partial F}{\partial N}\right)_{T,V} = -kT\ln\frac{Z_1}{N} = kT\ln\frac{N\Lambda^3}{V} = kT\ln\frac{P\Lambda^3}{kT}$$

• Internal energy  $U = -\frac{\partial}{\partial\beta} \log Z_N = \frac{3}{2} NkT$ 

• Entropy 
$$S = -\left(\frac{\partial F}{\partial T}\right)_{V,N} = Nk\left[\frac{5}{2} - \log\frac{N}{V}\Lambda^3\right]$$