

## Lecture spring 2024:

## General Relativity

### Problem sheet 7

↪ These problems are scheduled for discussion on **Thursday, 14 March 2024**.

#### Legend

\* If pressed for time, make sure to try solving the other problem(s) before attempting this one

#### **Problem 23**

In Minkowski space the d'Alembertian operator is given by  $\square = \eta^{\mu\nu} \partial_\mu \partial_\nu$ , which readily generalizes to the curved space-time (and coordinate-invariant) version  $\square = g^{\mu\nu} \nabla_\mu \nabla_\nu$ . Show that this latter expression is equivalent to

$$\square f = \frac{1}{\sqrt{-g}} \partial_\mu (\sqrt{-g} \nabla^\mu f) ,$$

a form which can be advantageous for calculations in practice.

#### **Problem 24**

Consider the stress-energy tensor for a perfect fluid,

$$T^{\mu\nu} = (\rho + p)u^\mu u^\nu + \eta^{\mu\nu} p$$

- a) Use the conservation of this tensor,  $\partial_\mu T^{\mu\nu} = 0$ , in the non-relativistic limit to derive familiar results from classical hydrodynamics. In particular, derive the

- continuity equation,  $\partial_t \rho + \nabla \cdot (\rho \mathbf{v}) = 0$
- Euler Equation,  $\partial_t \mathbf{v} + (\mathbf{v} \cdot \nabla) \mathbf{v} = -\frac{\nabla p}{\rho}$

[Hint: Have a look at 1.9 in the book. ;) But make sure that you really understand each step!]

- b)\* Now generalize the result to the Navier-Stokes equations, by allowing for off-diagonal (shear) terms in  $T_{\mu\nu}$ .

#### **Problem 25**

Noether's theorem states that for every *symmetry* in the Lagrangian – a field transformation that leaves the equations of motion invariant – there is a *conserved quantity*.

There thus exists a 4-vector  $j_N^\mu$  that is divergence-free,  $\partial_\mu j_N^\mu = 0$ , and the conserved quantity (the ‘Noether charge’) is given by  $Q_N = \int d^3x j_N^0$ . You are encouraged to look up the general proof (it’s one of the most important results in theoretical physics – and actually not that difficult!), which allows you to compute  $j_N^\mu$  for any such field transformation.

Let us now consider the specific case of space-time translations  $x^\nu \rightarrow x^\nu + a^\nu$ , where  $a^\mu$  is a constant 4-vector. Any theory formulated in Minkowski space must be invariant under these transformations, so there must be *four* conserved currents  $j_N^{(\nu)\mu} \equiv T^{\mu\nu}$  (one for each value of  $\nu$ ) and the corresponding charges are the total energy and 3-momentum of the system, respectively. The energy-momentum tensor, of any field  $\phi^a(x)$ , obtained in this way is given by

$$T^{\mu\nu} = -\frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi^a)} \partial^\nu \phi^a + \eta^{\mu\nu} \mathcal{L},$$

where  $\mathcal{L} = \mathcal{L}(\phi^a, \partial_\mu \phi^a)$  is the Lagrangian (density). Now consider the general Lagrangian of a single scalar field in an arbitrary potential,  $\mathcal{L} = -\frac{1}{2}(\partial_\mu \phi)^2 - \frac{1}{2}m^2 \phi^2$ , and derive the stress-energy tensor! Try to give an interpretation of the individual terms that appear in the resulting expression.