## FYS3520 - Problem set 8

## Spring term 2017

Corrections: Problem 2:  $E_n = 170$  corrected to 1700 keV

## Problem 1 – in class

- a) Figure 1 shows the electron spectra meassured on  $\beta$ -particles from decay of <sup>210</sup>Bi. What does this spectra tell about the nature of  $\beta$ -decay?
- b) Comment on the three decays in Figure 2. What can you say about the half-life of the three different decays?
- c) Show in a (semi-)classical calculation that the angular momentum transfer *l* (in units of  $\hbar$ ) to a electron in  $\beta$ -decay is much smaller then 1. Assume a typical *Q*-value of 3 MeV and *R*= 6 fm for the nuclear radius. What do we learn about possible and forbidden  $\beta$ -decays?
- d) Frank presents some cool things about his research



Figure 1: Electron spectra from  $\beta$ -decay of <sup>210</sup>Bi.



Figure 2:  $\beta$ -decays

## **Problem 2** $\beta$ -delayed neutron emission

After  $\beta$ -decay of a excited state it is not only possible to emit  $\gamma$ -rays. Depending on the excitation energy of the daughter nucleus it can also emit particles. The  $\beta$ -delayed neutron emission is essential for the control of nuclear reactors. However, in this example we analyze the  $\beta$ -delayed neutron emission of <sup>17</sup>O, which leads to the gs. of <sup>16</sup>O. Here <sup>17</sup>O has before been populated by  $\beta^-$ -decay of <sup>17</sup>N. The decay scheme is given in Figure 3

As given in the figure three neutrons are emitted with energies 383, 1171 and 1700 keV, which populate the ground state of  $^{16}$ O.

- a) Calculate the *Q*-value for the  $\beta$ -decay of <sup>17</sup>N.
- b) Derive a formula that allows you to calculate the excitation energy of <sup>17</sup>O as a function of the kinetic energy of the neutron and exicitation energy of <sup>16</sup>O. Calculate the excitation energy of <sup>17</sup>O.

Hint: You get the correct formula if you take into account energy *and* momentum conservation. You may regard only the <sup>17</sup>O-<sup>16</sup>O system and neglect <sup>17</sup>N.

c) Why can the  $\beta$ -delayed neutron emission not populate the first excited state of <sup>16</sup>O at 6.05 MeV?



Figure 3:  $\beta$ -delayed neutron emission from <sup>17</sup>O