# FYS 3610 Exercises week 39

## **Descriptive questions:**

- 1. Draw a sketch of the undisturbed Earth magnetic field using the Earth's rotational axis as a reference. Indicate direction of the magnetic field. What is the strength of the magnetic field near the equator and near the poles?
- 2. The Earth magnetic field are normally referred to the local coordinate systems (X,Y,Z) or (H, D, Z). Draw a figure that illustrates the Earth magnetic field vector decomposed in the two coordinate systems. Indicate geographic north in your figure.
- 3. Describe what happens when the solar wind, consisting of electrons and protons, hits the Earth magnetic field. Expalin the magnetopause currents.
- 4. What is a typical value for the solar wind stand-off distance? What are the controlling parameters for this distance; i.e. to push it inwards and outwards?
- 5. Explain the tail current.

### **EXERCISE 2**

Visit NASA's home page for the CGM model of the Earth's magnetic field: <u>http://omniweb.gsfc.nasa.gov/vitmo/cgm\_vitmo.html</u>

- i) Where is the north CGM pole located?
- ii) Calculate CGMLat, CGMLon, MLT, magnetic conjugate point, L- value, H, D, Z component for the magnetic field, magnetic field strength, Inclination for the following geographic co-ordinates: University of Oslo: (59.91, 10.73) University of Tromsø: (69.7, 18.9) Andøya Rocket Range: (69.28, 16.01) Longyearbyen: (78.2, 15.7)

#### **EXERCISE 3**

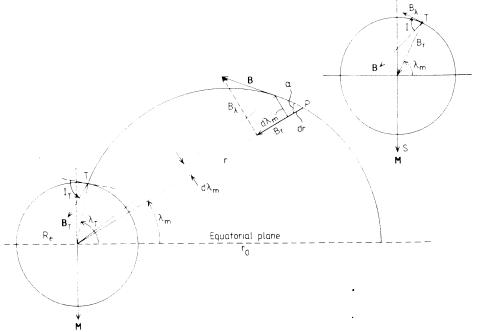


Figure 1: An illustration showing the geometry of the magnetic field line to assist in deriving a geometric formula for B.

Assume a dipole magnetic field. Introducing the magnetic latitude  $\lambda_m$  of which unit vector relates to the co-latitude unit vector defined in the lecture in the following way:

$$\widehat{\lambda}_m = -\widehat{\theta}$$

Then the magnetic field can be written as:

$$\vec{B} = B_r \hat{r} + B_{\lambda_m} = H_0 (-2\sin\lambda_m \hat{r} + \cos\lambda_m \hat{\lambda}_m)$$
  
i) Show that  $L = \frac{r_0}{R_E} = \frac{1}{\cos^2 \lambda_m}$   
Hint:  $\tan\alpha = \frac{r \cdot d\lambda_m}{dr} = \frac{B_\lambda}{B_r}$   
ii) Estimate B-field strength, inclination and L-val

ii) Estimate B-field strength, inclination and L-value based on the dipole model, and compare with corresponding values calculated in Exercise 2.

#### **EXERCISE 4**

Derive the equation for the standoff distance. Estimate the magnetopause standoff distance for average solar wind conditions and extreme conditions.