

# Week 37

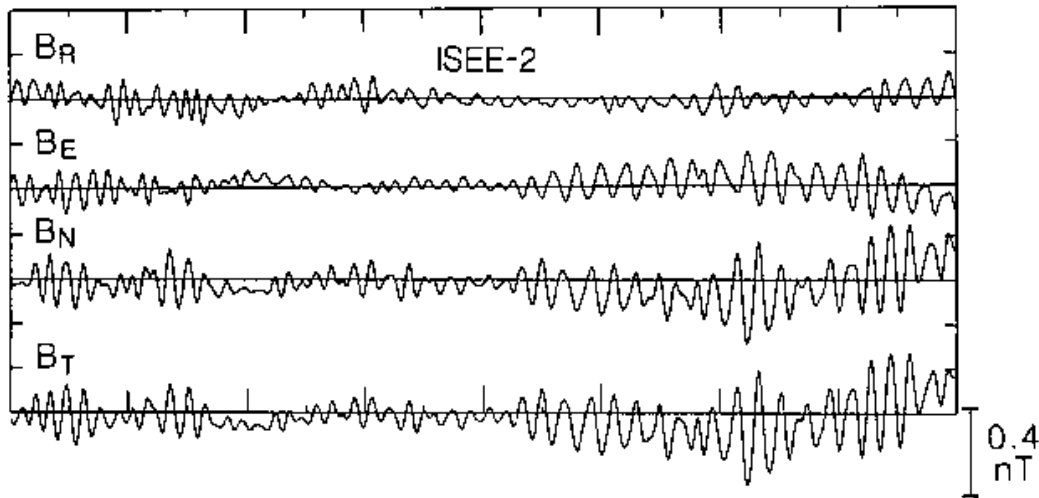
## Lecture 7:

- MHD : 14 equations/14 unknowns
- $R_m$ : Reynold magnetic number
- Ideal MHD : Frozen in flux

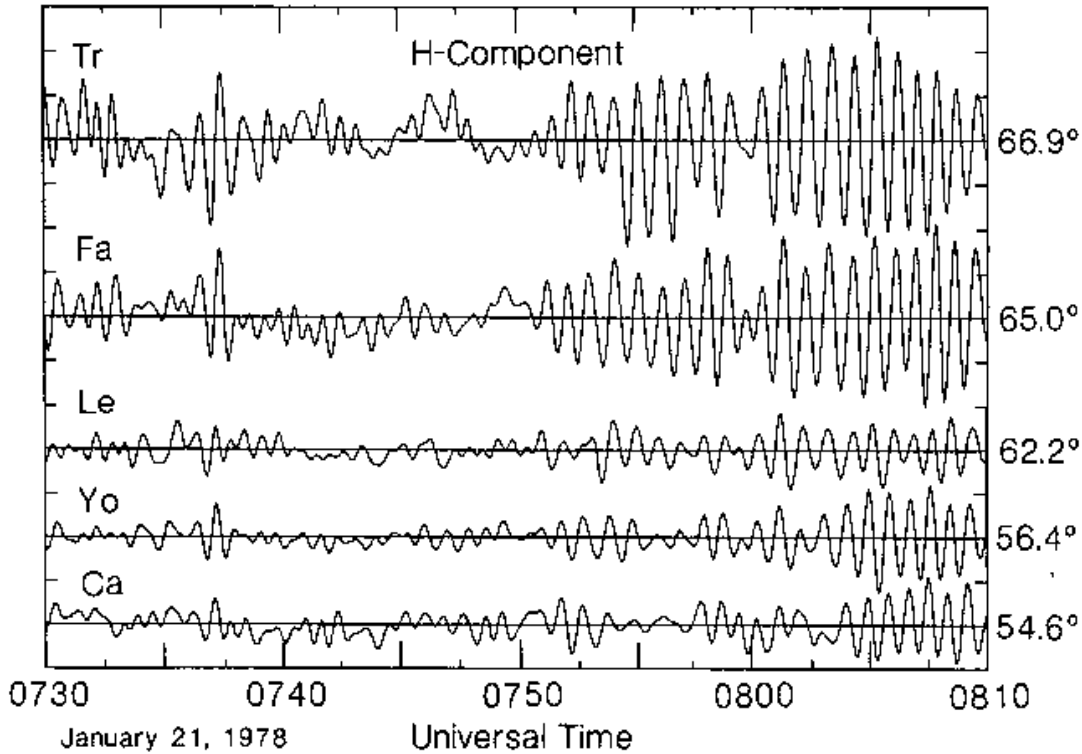
## Lecture 8:

MHD – waves:

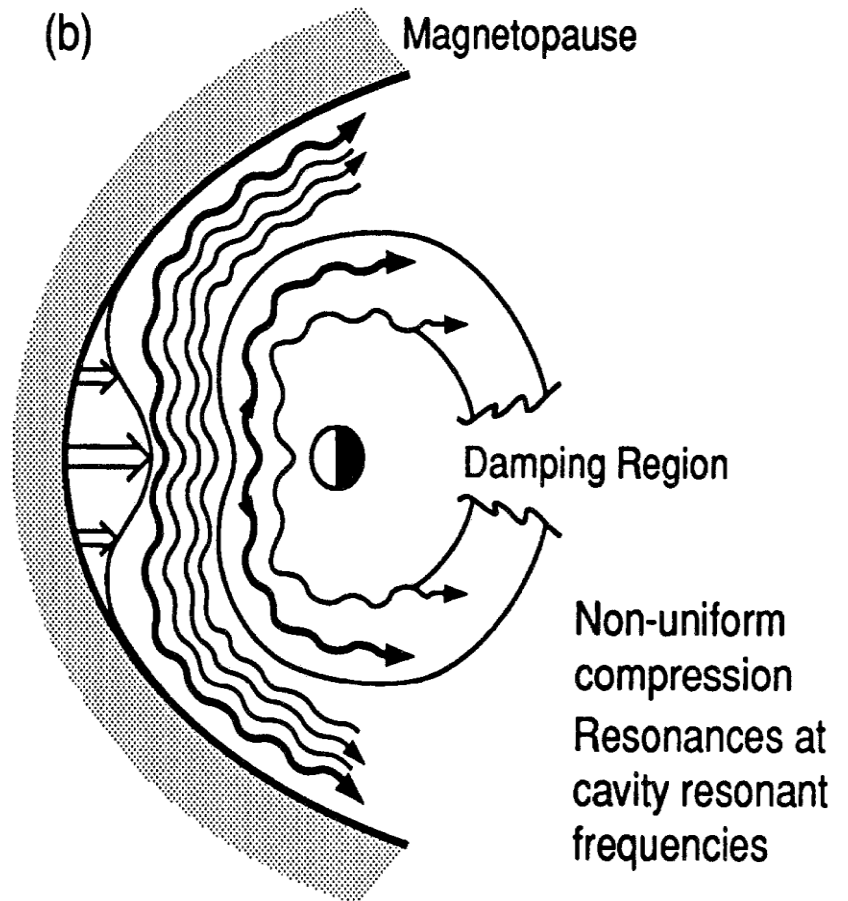
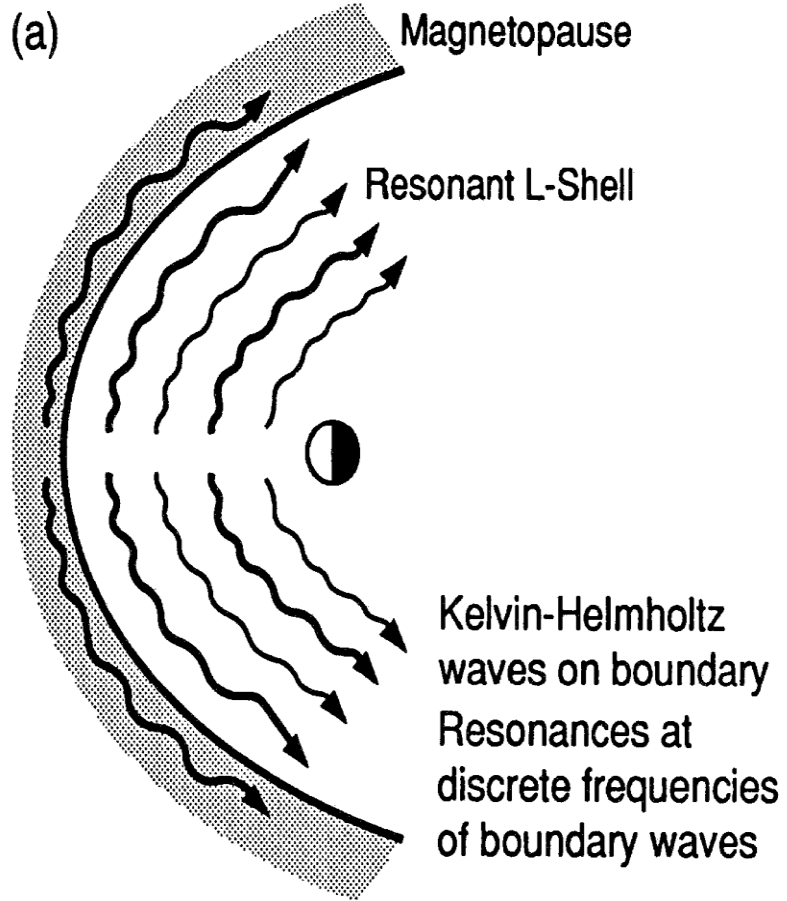
- Dispersion relation
- Plasma accoustic
- Alfvén
- Magnetosonic



Waves in solar wind

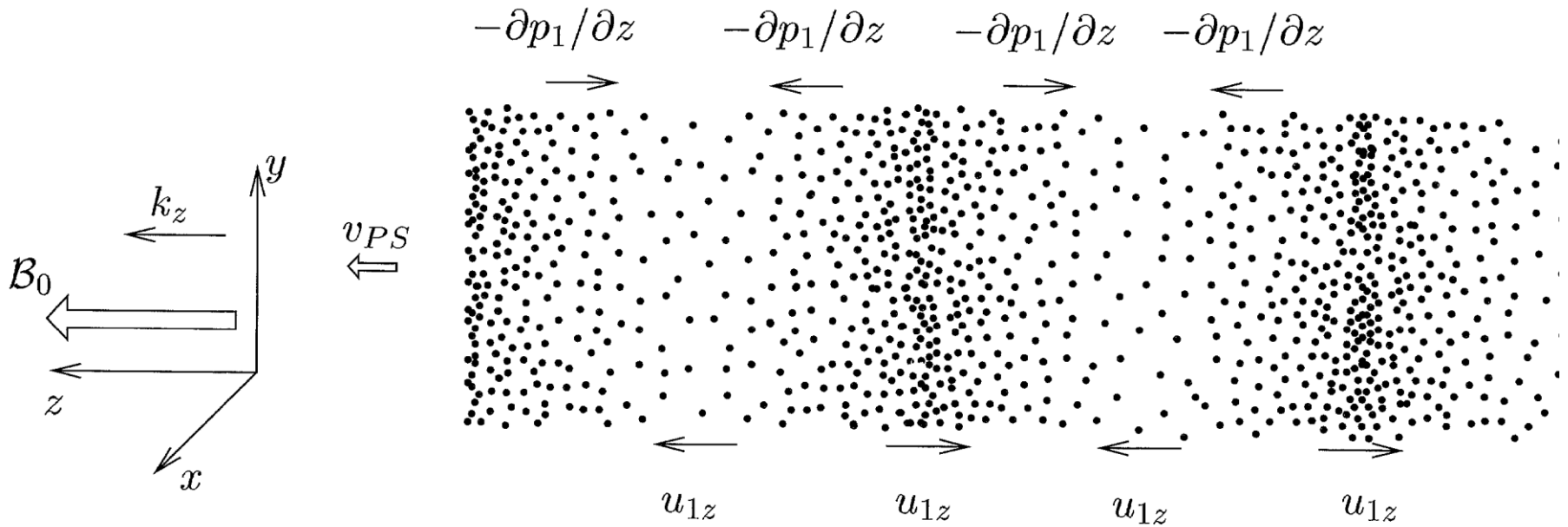


Corresponding waves in the Earth's magnetic field

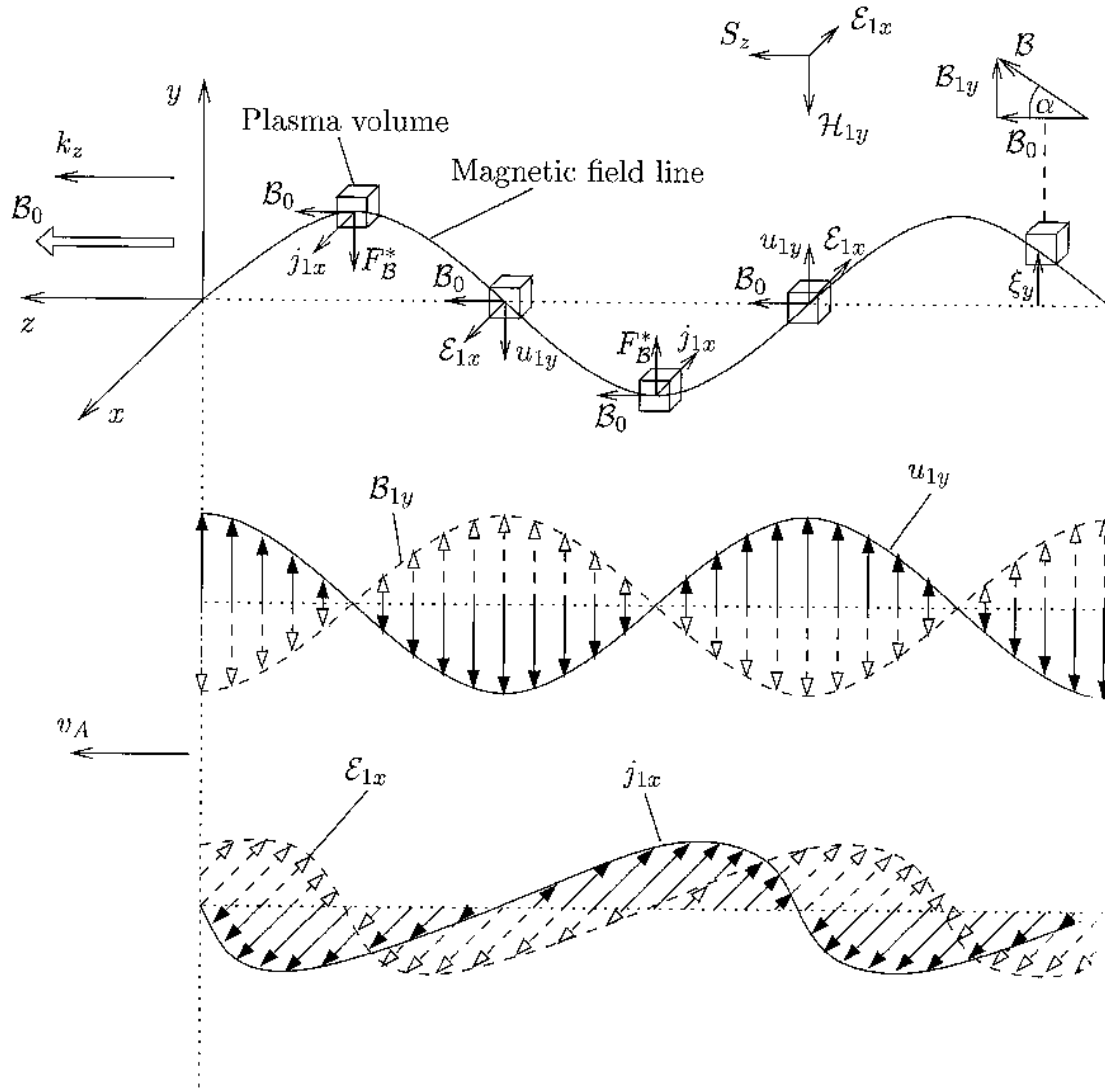


*(Note that this model explains peak amplitudes near noon)*

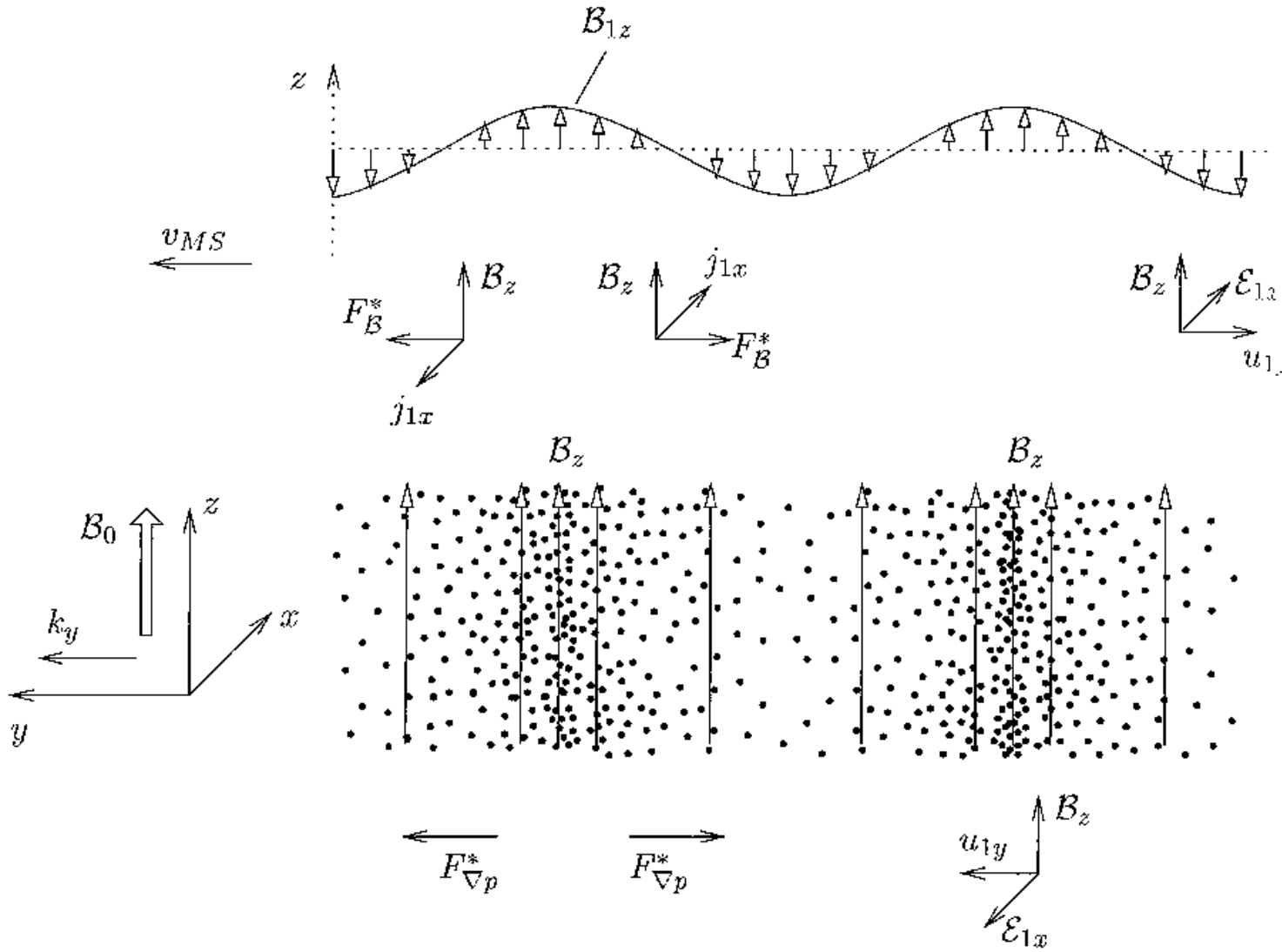
# Plasma Acoustic Waves



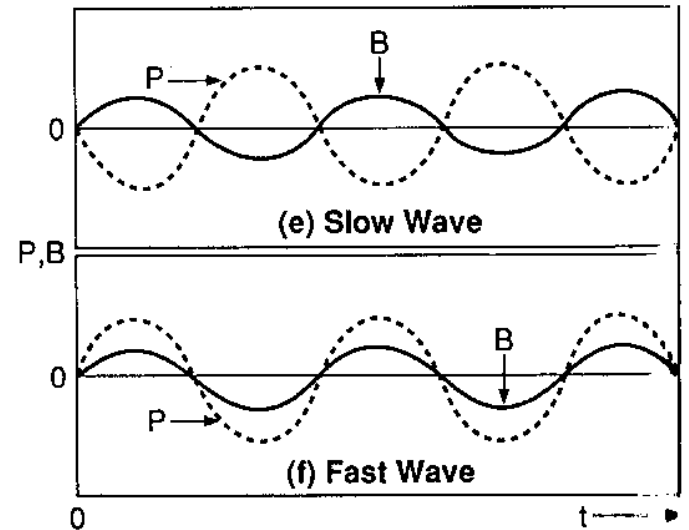
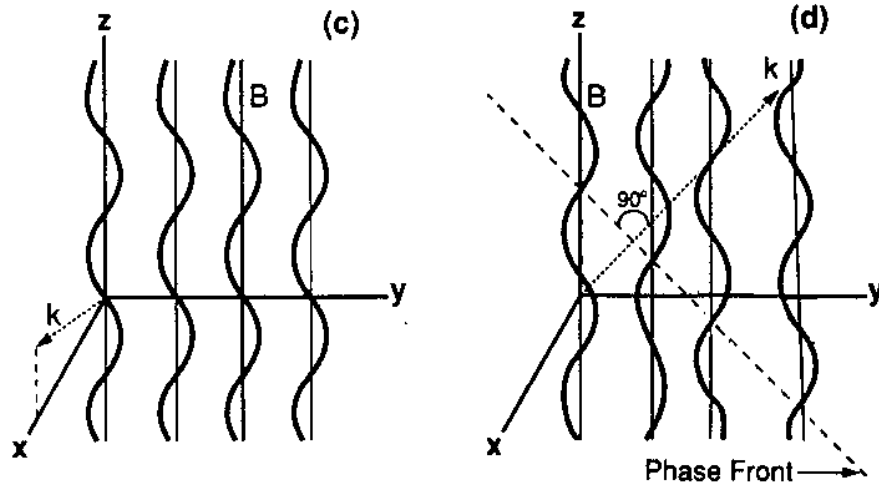
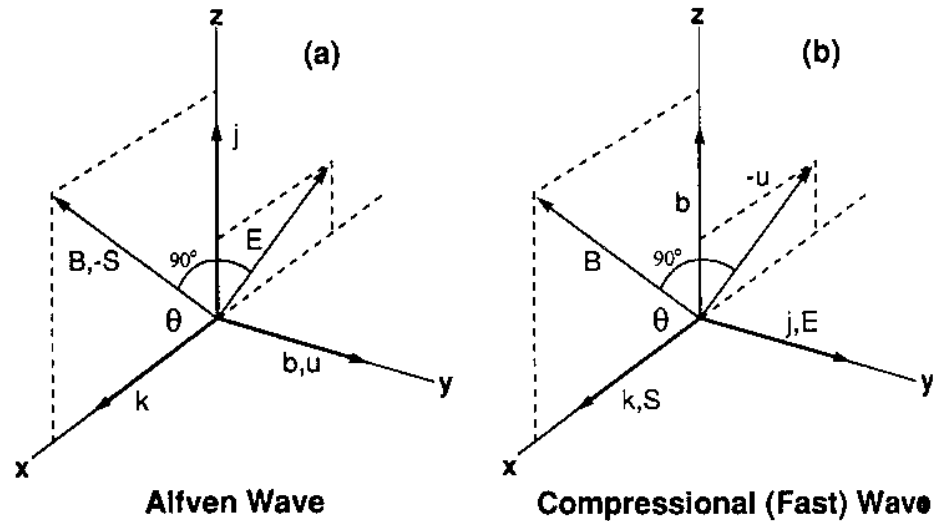
# Alfvén waves

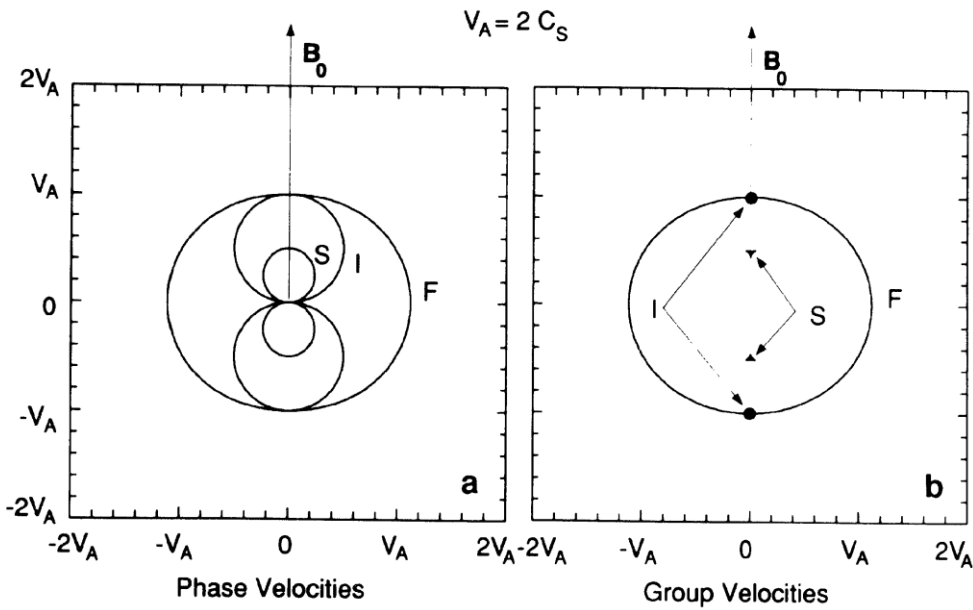


# Magnetosonic waves



**FIG. 11.2.** Schematic of wave polarizations for (a) the Alfvén wave and (b) the fast compressional wave. Displacements of the field lines (thick curves) at maximum displacement for (c) the Alfvén wave and (d) the fast compressional wave. The thin lines represent the unperturbed field. Plasma-pressure and magnetic-pressure perturbations versus time for (e) the slow compressional wave and (f) the fast compressional wave.

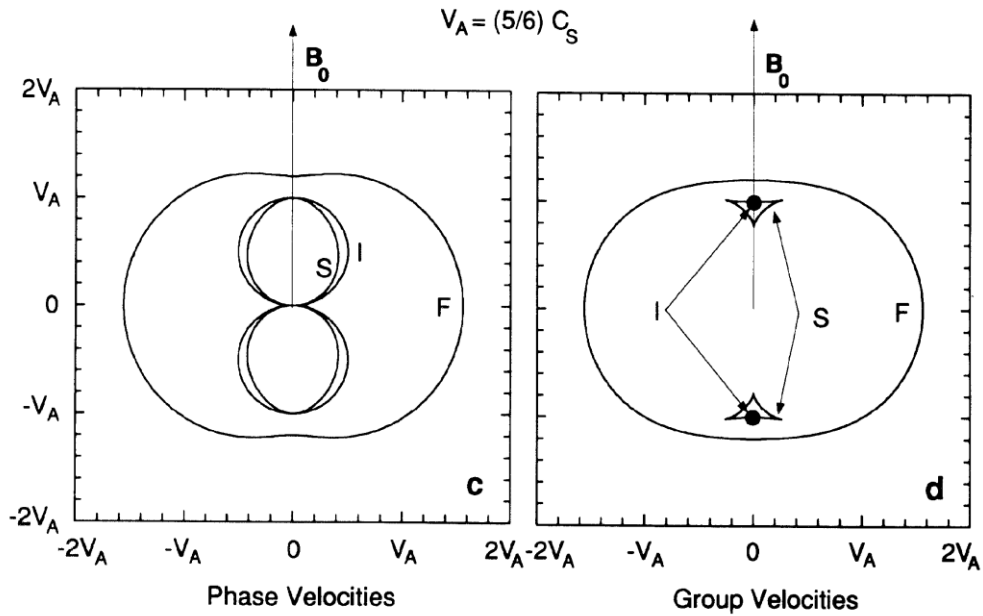




I = Alfvén wave

F(ast) and S(low)  
magnetosonic

$V_A$  = Alfvén velocity  
 $C_S$  = Speed of sound



Energy  
propagation