

Last time

Geodesics = "straight lines" on $S \subset \mathbb{R}^3$

$\Leftrightarrow \beta'(t) \& \underbrace{D\beta''(t)}$ parallel $\forall t$

\perp proj of

$\beta''(t)$ to $T_{\beta(t)}S$
 $\begin{matrix} \mathbb{R}^3 \\ \mathbb{R} \end{matrix}$

$\beta(t)$ constant speed geodesic $\Leftrightarrow D\beta''(t) = 0$

$\|\beta'(t)\| = \text{const} \in \mathbb{R}^2$
 local param $x = u \rightarrow s$

$\Leftrightarrow \beta(t) = x(u(t), v(t))$ sol. to the Geodesic equations

$$u'' + (u')^2 \Gamma_{11}^1 + 2u'v' \Gamma_{12}^1 + (v')^2 \Gamma_{22}^1 = 0$$

$$v'' + (u')^2 \Gamma_{11}^2 + 2u'v' \Gamma_{12}^2 + (v')^2 \Gamma_{22}^2 = 0$$

(*)

Γ_{ij}^k the Christoffel symbols def by

$$x_{uu} = \Gamma_{11}^1 x_u + \Gamma_{11}^2 x_v + eN$$

$$x_{uv} = \Gamma_{12}^1 x_u + \Gamma_{12}^2 x_v + fN$$

$$x_{vv} = \Gamma_{22}^1 x_u + \Gamma_{22}^2 x_v + gN$$

Rmk In general very difficult to solve (*)

explicitly.

But (*) is a second order ODE

(ordinary differential eqn)

Picard-Lindelöf Thm

\Rightarrow if we specify initial cond $u(0), v(0), u'(0), v'(0)$ it will

have a unique sol. defined on some time interval, and the sol. varies smoothly wr the initial conditions \Rightarrow

Prop 5.6.11 Every point of a Riemannian surface S

has an open nbhd V s.t.:

$\exists \epsilon, \tau > 0$ s.t. $\forall q \in V$ & $w \in T_q S$ w/ $\|w\| < \epsilon$,

there is a unique constant speed geodesic

$\gamma_w^q : (-\tau, \tau) \rightarrow S$ s.t. $\gamma_w^q(0) = q, (\gamma_w^q)'(0) = w$
initial conditions

Moreover, $\gamma_w^q(t)$ depends smoothly on q, w, t .

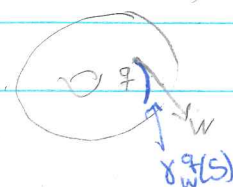
Rmk (*) homogeneous (does not depend explicitly on t)

\Rightarrow if $\gamma(t)$ is a sol, then $\eta(t) = \gamma(ct)$ is also a sol.

w/ initial cond $\eta(0) = \gamma(0), \eta'(0) = c\gamma'(0)$

$\Rightarrow \gamma_{cw}^q(t) = \gamma_w^q(ct)$ (**)

\Rightarrow may assume $c > 1$



~> The exponential map

Let $B_q(\epsilon) = \{W \in T_q S; \|W\| < \epsilon\}$, ϵ from Prop 5.6.11.

Def The map $\exp_q : B_q(\epsilon) \rightarrow S$, $\exp_q(W) = \gamma_W^q(1)$

is called the exponential map

Rmk $\exp_q(sW) = \gamma_{sW}^q(1) = \gamma_W^q(s)$ by (**)

$\Rightarrow s \mapsto \exp_q(sW)$ is a geodesic for each $W \in B_q(\epsilon)$

& $\frac{d}{ds} \exp_q(sW) \Big|_{s=0} = W$ (motivates the name)

These are the most important properties of the exponential map:

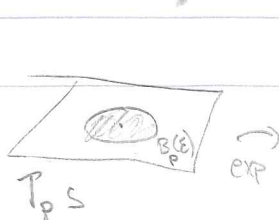
Thm 5.6.12 For every $p \in S$ there is an $\epsilon > 0$ s.t.

(1) \exp_p is a diffeo between $B_p(\epsilon)$ & a nbhd V of p
def a normal nbhd

(2) If ϵ is small enough, any 2 pts in

$\exp_p(B_p(\epsilon))$ can be joined by a unique

geodesic of length $< 2\epsilon$.



From this we get local parametrizations of $S \rightsquigarrow$

§5.7 Geodesic polar coordinates

Given $p \in S$, assume we have fixed an ON-basis

(w.r.t. \langle, \rangle_p) of $T_p S \rightsquigarrow$ cartesian coord (u, v) on $T_p S$

\rightsquigarrow polar coord (r, θ) on $T_p S$ s.t.

$(u, v) = (r \cos \theta, r \sin \theta) \Rightarrow$ get local param

$$x: \underbrace{B_p(\epsilon)}_{T_p S} \rightarrow S, \quad x(r, \theta) = \exp_p(r \cos \theta, r \sin \theta),$$

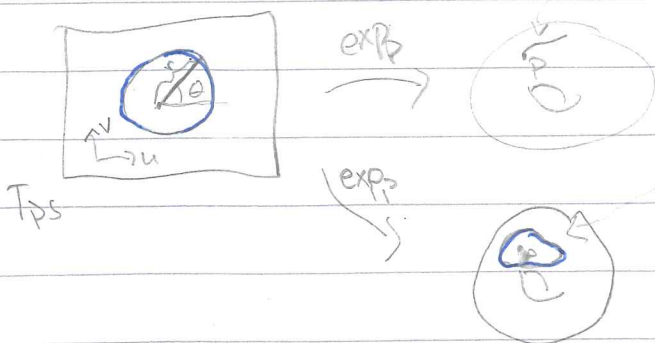
$(r, \theta) \in (-\epsilon, \epsilon) \times \underbrace{J}_{\text{interval of length } < 2\pi}$
(to get 1-1)

Geodesic polar coordinates

Def $r = \text{const} \Rightarrow \exp_p((r \cos \theta, r \sin \theta))$ geodesic circle

$\theta = \text{const} \Rightarrow$

geodesic radii



the coordinate

curves for

geodesic polar coord.

△ Geodesic radii are geodesics,

△ geodesic circles are usually not

Rmk Geodesic radii constant speed geodesics,

w.l.o.g assume they are param by arc-length,

⇒ for the metric $ds^2 = E(r, \theta) dr^2 + F(r, \theta) dr d\theta + G(r, \theta) d\theta^2$

we always have $E(r, \theta) = 1$

$\langle x_r, x_r \rangle$

$$E(r, \theta) = 1$$

Lma 5.7.2 (Gauss' lma) $F(r, \theta) = 0$ for (r, θ)
geodesic polar coord.

Pf $Dx_{rr} = 0$ since $r \mapsto x(r, \theta)$ wr θ fixed are
constant speed geodesics.

$$\text{But } Dx_{rr} = \Gamma_{rr}^1 x_r + \Gamma_{rr}^2 x_\theta \Rightarrow \Gamma_{rr}^1 = \Gamma_{rr}^2 = 0$$

$$\Rightarrow 0 = \Gamma_{rr}^1 F + \Gamma_{rr}^2 G = F_r - \frac{1}{2} E_\theta$$

↑
from formula for the Christoffel symbols

$$\text{But } E \equiv 1 \Rightarrow E_\theta = 0 \Rightarrow F_r = 0$$

$$\Rightarrow F(r, \theta) = F(0, \theta) = \langle x_r(0, \theta), x_\theta(0, \theta) \rangle = 0$$

$= 0$ □

Ex 5.7.1 & 5.7.3 Geodesic polar coord & metric for $\mathbb{E}^3, S^2, \mathbb{D}$.

(1) \mathbb{E}^2 : At $p=0$: $x(r, \theta) = (r \cos \theta, r \sin \theta)$

Metric $G = \langle x_\theta, x_\theta \rangle = r^2 \Rightarrow ds^2 = dr^2 + r^2 d\theta^2$

2) S^2 : At $p = (0, 0, 1)$

Know Geodesics are given by $\gamma(s) = A \cos(s) + B \sin(s)$

w. initial cond. $\gamma(0) = A, \gamma'(0) = B$

$\Rightarrow A = (0, 0, 1), \gamma(s) = \exp_p(sB)$

and since $T_p S \cong \mathbb{R}^2 \times \{0\} = \{(w_1, w_2, 0) \in \mathbb{R}^3\} \cong \mathbb{R}^2$

we get $\exp_p(r(\cos \theta, \sin \theta)) = (0, 0, 1) \cos r + (\cos \theta, \sin \theta, 0) \sin r$

$\Rightarrow x(r, \theta) = (\sin r \cos \theta, \sin r \sin \theta, \cos r)$

Metric $G = \langle x_\theta, x_\theta \rangle = \langle (-\sin r \sin \theta, \sin r \cos \theta, 0), (-\sin r \sin \theta, \sin r \cos \theta, 0) \rangle$

$= \sin^2 r \sin^2 \theta + \sin^2 r \cos^2 \theta = \sin^2 r$

$\Rightarrow ds^2 = dr^2 + \sin^2 r d\theta^2$

① : $p=0$. know geodesics are radii

Q Why?

$\Rightarrow x(r, \theta) = (h(r)\cos\theta, h(r)\sin\theta)$ for some $h(r)$

Some different ways of finding $h(r)$

1) $r \mapsto (\underbrace{h(r)\cos\theta}_u, \underbrace{h(r)\sin\theta}_v)$ should be param.

by arc-length \Rightarrow

$$E(u,v)\left(\frac{du}{dr}\right)^2 + F(u,v)\frac{du}{dr}\frac{dv}{dr} + G(u,v)\left(\frac{dv}{dr}\right)^2 = 1$$

$$= \frac{4(h'(r))^2\cos^2\theta + 0 + 4(h'(r))^2\sin^2\theta}{(1-u^2-v^2)^2}$$

$$\Rightarrow 4(h'(r))^2 = (1-h^2(r))^2 \quad \text{solve this}$$

Hint $\tanh x$ satisfies $1 - \tanh^2 x = (\tanh' x)'$

2) Solve the geodesics eqn in \mathbb{D} under the assumption

that a sol is given by $(u(r), 0)$ + use rotation symmetry.

3) $r \mapsto (h(r)\cos\theta, h(r)\sin\theta)$ should be param. by

arc-length \Rightarrow the hyperbolic length of $(h(r)\cos\theta, h(r)\sin\theta)$ should be given by r

and we recall from what we did before that

a ray from the origin in \mathbb{D} of hyperbolic length

r has euclidean length $\tanh \frac{r}{2} \Rightarrow$

$$h(r) = \tanh \frac{r}{2}.$$

Metric $x_\theta = (\underbrace{-\tanh \frac{r}{2} \sin \theta}_a, \underbrace{\tanh \frac{r}{2} \cos \theta}_b)$

$$\Rightarrow \langle x_\theta, x_\theta \rangle = 4 \frac{a^2 + b^2}{(1 - u^2 - v^2)^2} = \frac{4 \tanh^2(r/2)}{(1 - \tanh^2(r/2))^2} = \sinh^2(r)$$

$$\Rightarrow ds^2 = dr^2 + \sinh^2 r d\theta^2$$