Case studies

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- Already considered the SST data several times
- On interannual time scales, the dominant features in the process are the so-called El Niño and La Niña phenomena
- Cressie and Wikle (2011):

due to its complexity and uncertainties related to the fundamental physical mechanisms that drive these phenomena, SST is one of the few processes in oceanography in which "statistica" long-lead forecast models perform as well as or better than deterministic models

- Now: Modelling through reduced-dimension
- Both linear and "nonlinear" dynamics
- Aim: 6 month ahead forecast

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SST



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Model

 $\begin{aligned} \mathbf{Z}_{t} = \mathbf{Y}_{t} + \boldsymbol{\nu}_{t}, & \boldsymbol{\nu}_{t} \stackrel{ind}{\sim} \mathcal{N}(\mathbf{0}, \sigma_{\nu}^{2} \mathbf{I}) \\ \mathbf{Y}_{t} = \mathbf{\Phi} \boldsymbol{\alpha}_{t} + \boldsymbol{\gamma}_{t}, & \boldsymbol{\gamma}_{t} \stackrel{ind}{\sim} \mathcal{N}(\mathbf{0}, \sigma_{\gamma}^{2} \mathbf{I}) \\ \boldsymbol{\alpha}_{t} = \mathbf{M} \boldsymbol{\alpha}_{t-1} + \boldsymbol{\eta}_{t}, & \boldsymbol{\nu}_{t} \stackrel{ind}{\sim} \mathcal{N}(\mathbf{0}, \boldsymbol{\Sigma}_{\eta}) \end{aligned}$

Of interest: $\mathbf{Y}_{t+\tau}^{\rho} = \mathbf{\Phi} \boldsymbol{\alpha}_{t+\tau}$. Note: $E[\mathbf{\Phi} \mathbf{\Phi}^T \mathbf{Z}_{t+\tau} | \boldsymbol{\alpha}] = \mathbf{\Phi} \mathbf{\Phi}^T \mathbf{\Phi} \boldsymbol{\alpha}_{t+\tau} = \mathbf{\Phi} \boldsymbol{\alpha}_{t+\tau} = \mathbf{Y}_{t+\tau}^{\rho}$ Inference through MCMC









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$$\begin{aligned} \mathbf{Z}_t = \mathbf{Y}_t + \boldsymbol{\nu}_t, & \boldsymbol{\nu}_t \stackrel{\text{ind}}{\sim} N(\mathbf{0}, \sigma_{\nu}^2 \mathbf{I}) \\ \mathbf{Y}_t = \mathbf{\Phi} \boldsymbol{\alpha}_t + \boldsymbol{\gamma}_t, & \boldsymbol{\gamma}_t \stackrel{\text{ind}}{\sim} N(\mathbf{0}, \sigma_{\gamma}^2 \mathbf{I}) \\ \boldsymbol{\alpha}_t = \mathbf{M} \boldsymbol{\alpha}_{t-1} + \boldsymbol{\eta}_t, & \boldsymbol{\nu}_t \stackrel{\text{ind}}{\sim} N(\mathbf{0}, \boldsymbol{\Sigma}_{\eta}) \end{aligned}$$

$$\begin{split} & \mathsf{vec}(\mathbf{M}) \sim & \mathsf{N}(\mathsf{vec}(0.9\mathbf{I}), 100\mathbf{I}) \\ & \mathbf{\Sigma}_{\eta}^{-1} \sim & \mathsf{Wishart}(\frac{1}{p-1}(100\mathbf{I})^{-1}, p-1) \\ & \sigma_{\nu}^{2} \sim & \mathsf{Inverse \ Gamma}(0.1, 100) \\ & \sigma_{\gamma}^{2} \sim & \mathsf{Inverse \ Gamma}(0.1, 100) \end{split}$$

- Fix $\alpha_t = \mathbf{\Phi}^T \mathbf{Z}_t$, t = 1, ..., 6With T = 349 it should not be very sensitive to this initial assumption
- Inference through MCMC using $p_{\alpha} = 10$.

Linear model - True and predicted processes



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Figure 9.3. Reduced-dimension linear DSTM patienties forecast density of monthly Nido 1.4 regions (05 5-5 N; 10.2⁻¹) (70 W) (SST averages. The top point (of aboves the positive forecast adneys) for October 1997 given data up to April 1997. The bottom panel (b) shows the corresponding posterior forecast density for October 1998 given data up to April 1996. In each panel, the hair "start" shows the means of the forecast distribution, and the red "star" shows the corresponding observed value (6 months later) of the Niko 3.4 region. Want different dynamics for the El Niño and La Niña phenomena Assume now \mathbf{M}_t changes with time:

$$\boldsymbol{\alpha}_t = \boldsymbol{\mu}_t + \boldsymbol{\mathsf{M}}_t \boldsymbol{\alpha}_{t-1} + \boldsymbol{\eta}_t, \qquad \qquad \boldsymbol{\eta}_t \stackrel{\scriptscriptstyle{\mathsf{ind}}}{\sim} \mathcal{N}(\mathbf{0}, \boldsymbol{\Sigma}_\eta)$$

Assume μ_t and M_t depend on which *climate regime* we are in:

$$\boldsymbol{\mu}_t = \boldsymbol{\mu}(I_t), \quad \mathbf{M}_t = \mathbf{M}(I_t)$$

Also extend model to

$$\mathbf{Z}_t = \mathbf{Y}_t + \boldsymbol{\nu}_t, \qquad \qquad \boldsymbol{\nu}_t \stackrel{ind}{\sim} N(\mathbf{0}, \boldsymbol{\Sigma}_{\nu})$$

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For current/previous time points:

$$I_t = \begin{cases} 0 & \text{if } SOI_t < \text{lower threshold} \\ 1 & \text{if } SOI_t \text{ in between} \\ 2 & \text{if } SOI_t > \text{upper threshold} \end{cases}$$

where SOI_t is the Southern Oscilation index. For future time points, we *model* I_t as a stochastic process:

$$I_t = \begin{cases} 0 & \text{if } W_t < \text{lower threshold} \\ 1 & \text{if } W_t \text{ in between} \\ 2 & \text{if } W_t > \text{upper threshold} \end{cases}$$

where we assume

$$\begin{split} \mathcal{N}_t | \boldsymbol{\beta}_w, \sigma_w^2 \sim & \mathsf{Gau}(\mathbf{X}_t^T \boldsymbol{\beta}, \sigma_w) \\ \mathbf{X}_t \equiv & (1, U_t, U_t \sin(2\pi b_t/12), U_t \cos(2\pi b_t/12), U_t^2)^T \end{split}$$

where U_t is an east-west component of wind at 10 meters above surface and b_t is an index of month.

$$\boldsymbol{\alpha}_{t} = \begin{pmatrix} \boldsymbol{\alpha}^{(1)} \\ \boldsymbol{\alpha}^{(2)} \end{pmatrix}, \quad \boldsymbol{\mu}_{t} = \begin{pmatrix} \boldsymbol{\mu}^{(1)}(\boldsymbol{l}_{t}) \\ \boldsymbol{0} \end{pmatrix}, \quad \boldsymbol{\mathsf{M}}_{t} = \begin{pmatrix} \boldsymbol{\mathsf{M}}^{(1,1)}(\boldsymbol{l}_{t}) & \boldsymbol{\mathsf{M}}^{(1,2)} \\ \boldsymbol{\mathsf{M}}^{(2,1)} & \boldsymbol{\mathsf{M}}^{(2,2)} \end{pmatrix}$$

- Fitting performed by MCMC
- We can get posterior probabilities for climate regimes. This can be used for prediction:

$$E[\boldsymbol{\alpha}_{t+1}|\mathbf{Z}_{1:t}] = \sum_{j=0}^{2} \Pr(I_{t+1} = j|\mathbf{Z}_{1:t}) E[\boldsymbol{\mu}(j) + \mathbf{M}(j)\boldsymbol{\alpha}_{t}|\mathbf{Z}_{1:t}]$$

Nonlinear model - Posterior forecast of montly Niño region



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Nonlinear model - Posterior forecast of montly Niño region



Figure 94. Reduced-dimension mediance DSTM potenties forecast density of monthly Nido 3.4 region ($5^{-5} - 5^{-1}$, $10^{-2} - 10^{-2}$ W) STS reverges. The top point is above the distribution for Checher 1997, ziven data up to April 1997. The bottom part is above the point of the stars for Checher 1998, given data up to April 1998. In each part, the Man 'near' above the mean of the forecast distribution and the red 'star" shows the corresponding observed wite (5 month) larger of the Nido 3.4 region.

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Regime-dependent predictions



Modelling and forecasting the Eurasian Collared Dove invasion (ECD)

- Ecologist interested in growth and dispersal of biotic organisms.
- Dynamics particularly important
- ECD: Counts 1986 2002 at peak of the avian breeding season
- Survey routes of approximately 39.4 km
- 50 stops, observer counts and identifies by sight and sound for a 3-minutes period



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Assume $Y_t(\mathbf{s}_i)$ is the true (unknown) number of ECD counts. Assume

 $Z_t(\mathbf{s}_i) | \mathbf{Y}_t(\mathbf{s}_i), \theta \stackrel{ind}{\sim} \text{Binomial}(Y_t(\mathbf{s}_i), \theta)$

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- Information on θ possible through *replicable* samples. Not available

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- Replicate counts were taken on a different, but related species. Can be used to construct an *informative* prior for θ: θ ~ Beta(a_θ, b_θ)

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- Replicate counts were taken on a different, but related species. Can be used to construct an *informative* prior for θ: θ ~ Beta(a_θ, b_θ)
- Gives (assuming different θ 's for each observation)

$$[Z_t(\mathbf{s}_i)|\mathbf{Y}_t(\mathbf{s}_i)] = \int_0^1 [Z_t(\mathbf{s}_i)|\mathbf{Y}_t(\mathbf{s}_i),\theta][\theta] d\theta = \mathsf{Beta-Bin}(Y_t(\mathbf{s}),a_\theta,b_\theta)$$



Numbers at observation locations

$$\mathbf{Y}_t = (Y_t(\mathbf{s}_1), ..., Y_t(\mathbf{s}_m))^T$$

Intensity at observation and prediction locations

$$\boldsymbol{\lambda}_t = (\lambda_t(\mathbf{s}_1), ..., \lambda_t(\mathbf{s}_m), \lambda_t(\mathbf{s}_{m+1}), ..., \lambda_t(\mathbf{s}_n))^T$$

Assume

$$\begin{split} Y_t(\mathbf{s}_i) | \boldsymbol{\lambda}_t & \stackrel{ind}{\sim} \mathsf{Poisson}(\lambda_t(\mathbf{s}_i)), i = 1, ..., n \\ \mathbf{Y}_t | \boldsymbol{\lambda}_t & \stackrel{ind}{\sim} \mathsf{Poisson}(\mathbf{H}\boldsymbol{\lambda}_t) \end{split}$$

H is an incidence matrix

$$\begin{split} \boldsymbol{\lambda}_t = & \mathbf{M} \boldsymbol{\lambda}_{t-1} \\ = & \mathbf{B}(\boldsymbol{\tau}) \mathbf{G}(\boldsymbol{\lambda}_{t-1}; \boldsymbol{\theta}^G) \boldsymbol{\lambda}_{t-1} \\ & \mathbf{G}(\boldsymbol{\lambda}_{t-1}; \boldsymbol{\theta}^G) : \text{Diagonal matrix for growth over time} \\ & G_{ii} = \exp\left\{\theta_1^G \left(1 - \frac{\lambda_t(\mathbf{s}_i)}{\theta_2^G}\right)\right\} \\ & \mathbf{B}(\boldsymbol{\tau}) : \text{Dispersal} \\ & B_{ij} \propto \exp\left\{-\frac{-d_{ij}^2}{\tau(\mathbf{s}_i)}\right\}, \quad \sum_j B_{ij} = 1 \end{split}$$

- G: Ricker growth model
- Only random parts: $oldsymbol{\lambda}_1$ and priors on parameters $(oldsymbol{ heta}^{\scriptscriptstyle G},oldsymbol{ au})$
- Corresponds to a discretized diffusion equation with a population that can pass through "open boundaries"

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- λ_1 is a Gaussian spatial process with exponential covariance function at log scale
- Similar for au
- Allow for predictions at points $\mathbf{s}_{m+1}, ..., \mathbf{s}_n$.
- Dynamical model allows or forcasting

Dove: Spatial predictions and forecast



Figure 9.11 Posterior means of ECD abundance in the United States for the period 1980 (corresponding to the sampling period considered in this analysis). Areas of the map s "white" were not included in the prediction grid.



Figure 9.12 Posterior mean of the ECD-abundance forecast in the United States for th 2003 to 2020 (where the years 2004–2020 correspond to out-of-sample forecasts). Areas of shown as "white" were not included in the prediction grid.

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Dove: Specific locations



Figure 9.13 Comparisons of posterior credible intervals for the ECD population from 1986 to 2020. Top panel: Credible intervals for a location in south Florida (red) and for eastern Kansas (black). Bottom panel: Credible intervals for the south Florida location (red) and for northern Utah (black).

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Figure 9.14 Posterior mean of the spatially varying dispersal coefficients (τ) in the ECD process nodel (9.19).

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- Implementation through MCMC
- Can also obtain uncertainties
 - Look similar to posterior mean (variance equal to expectation in Poisson distribution)
- Posterior distributions of growth and dispersal parameters where more narrow than their priors, indicating learning from data
- Spatially varying au's was supported by data